

# 8.022 (E&M) – Lecture 11

## Topics:

- Introduction to Special Relativity
  - Length contraction and Time dilation
  - Lorentz transformations
  - Velocity transformation

## Special relativity

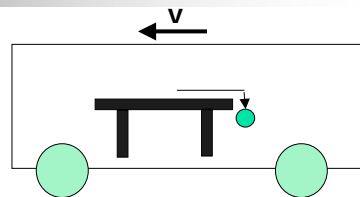
- Ready for the challenge?
  - Special relativity seems easy but it's not!
  - A new way of thinking that often goes against intuition
  - It will take some time to "digest it", but believe me: it's worth the effort!
- Why do we need it in 8.022?
  - Weren't you frustrated last time when magnetic forces came out of nowhere?
  - Special relativity naturally explains them in terms of electric forces seen from in a reference frame in motion
- This is important for everybody
  - Physics majors: first of many iterations on a crucial topic
  - Non Physics majors: chance to know what you are missing
    - Don't forget: you are still in time...

# The principles of special relativity

- Formulated in 1905 by A. Einstein
  - Incredible but true:  
no Nobel Prize for this!
- Based upon 2 postulates
  - The laws of physics are the same for all reference frames
  - The speed of light is the same ( $c$ ) in all reference frames
- (Inertial) Reference frame
  - System of coordinates in which the observer is non accelerating (inertial = non accelerating)

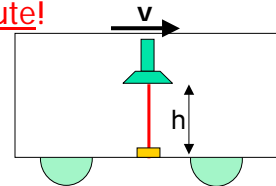
# Reference frames: examples

- Situation
  - A train is moving with velocity  $v$  w.r.t. to a station
  - A table is anchored to the train
  - A ball is falling from the table
- We can identify 3 systems of reference and 3 observers:
  - Observer 1: sitting on a bench at the station
  - Observer 2: sitting on the table on the train
  - Observer 3: a bug sitting on top of the falling ball
- Who are the observers in an inertial reference frame?
  - Observers 1 and 2
  - Observer 3 is not: the ball is falling with acceleration  $g$



## Is time the same in all reference frames?

- These (apparently) innocent assumptions have amazing consequences such as time is not absolute!
- Problem**
  - The train is moving with velocity  $v$  // x axis
    - Observer 1: standing in the train
    - Observer 2: at the station
  - Observer 1 flashes a pulse of light vertically to a photosensor mounted on the floor of the train
  - Both observers measure the time between when the light is emitted and when the light reaches the sensor



Will the 2 observers measure the same time?

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## Time in different reference frames

- Let's calculate time measured by the 2 observers
- Train reference frame (observer 1)

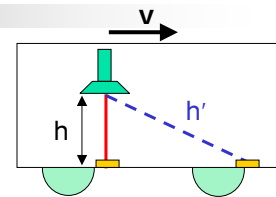
$$\begin{cases} \text{Distance traveled by light: } h \\ \text{Velocity of light: } c \end{cases} \Rightarrow \Delta t = \Delta t_1 = \frac{h}{c}$$

- Station reference frame (observer 2)

$$\begin{cases} \text{Distance traveled by light: } h' = \sqrt{h^2 + (v\Delta t_2)^2} \\ \text{Velocity of light: } c \end{cases} \Rightarrow \Delta t' = \Delta t_2 = \frac{h'}{c}$$

$$(\Delta t_2)^2 = \left(\frac{h'}{c}\right)^2 = \frac{h^2 + (v\Delta t_2)^2}{c^2} = \Delta t_2^2 + \frac{v^2}{c^2} \Delta t_2^2 \Rightarrow \Delta t_1 = \Delta t_2 \sqrt{1 - \frac{v^2}{c^2}}$$

$$\text{Defining } \gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \Rightarrow \boxed{\Delta t' = \gamma \Delta t}$$



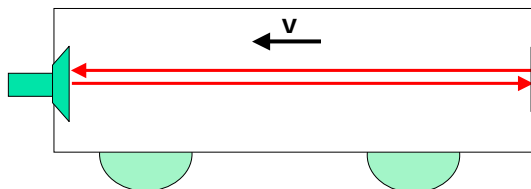
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## Time dilation

- We just derived a very important result!
- Gamma factor:  $\gamma \equiv \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{1}{\sqrt{1 - \beta^2}} > 1$  with  $\beta \equiv \beta_v \equiv \frac{v}{c}$
- Since  $\Delta t' = \gamma \Delta t \rightarrow \Delta t'$  is always larger than  $\Delta t$ 
  - $\Delta t'$  = time measured by the observer in the station who sees the clock in motion
  - $\Delta t$  = time measured by the observer on the train, at rest wrt the clock
- Conclusion:  
Clocks in motion run slower (time dilation)  $\Delta t' = \gamma \Delta t$

## Length in different reference frames

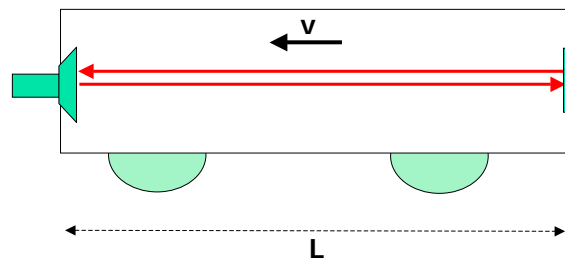
- Problem 2
  - Now observer 1 flashes a pulse of light horizontally from left end of the train
  - The light is reflected by a mirror on the right end wall and detected by a photosensor on the left wall



What is the length of the train measured by each observer?

## Length in train reference frames

- For observer in train reference frame
  - Events we are interested in: emission and reception of light
    - Time in between the two:  $\Delta t = \Delta t_{\text{train}}$
    - Length of the train:  $L = \frac{c \Delta t}{2} \Rightarrow \Delta t = \frac{2L}{c}$



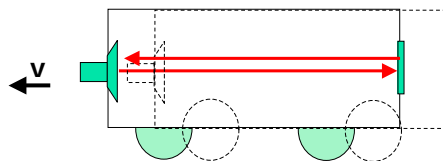
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## Length in the station reference frame

- Calculate separately  $\Delta x_1$  (L→R) and  $\Delta x_2$  (R→L)



$$\Delta t'_1 = (L' - v \Delta t'_1) / c$$

$$\Delta t'_2 = (L' + v \Delta t'_2) / c$$

- $\Delta t_1$  is shorter because train and light move in opposite directions
- $\Delta t_2$  is longer because train and light move in the same direction
- $L' (t')$  = length (time) measured from station reference frame

- Rearrange terms:

$$\begin{cases} \Delta t'_1 = \frac{L'}{c + v} \\ \Delta t'_2 = \frac{L'}{c - v} \end{cases}$$

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## Length contraction

- Total time in the station reference frame = sum of  $\Delta t'_1$  and  $\Delta t'_2$ :

$$\begin{aligned}\Delta t' &= \Delta t'_1 + \Delta t'_2 = \frac{L'}{c-v} + \frac{L'}{c+v} = \\ &= L' \frac{2c}{c^2 - v^2} = L' \frac{2c}{c^2(1 - \frac{v^2}{c^2})} = \frac{2L'\gamma^2}{c}\end{aligned}$$

- Remember how time dilates:  $\Delta t' = \gamma \Delta t \rightarrow$

$$\frac{2L'\gamma^2}{c} = \Delta t' = \gamma \Delta t = \gamma \frac{2L}{c} \Rightarrow \boxed{L' = \frac{L}{\gamma}}$$

- Since  $\gamma > 1 \rightarrow$

Moving objects appear contracted (length contraction)

## Summary so far

- Assume Special Relativity postulates hold:

- The laws of physics are the same for all reference frames
- The speed of light is the same ( $c$ ) in all reference frames

- Consequences:

- Time dilation
  - clocks in motion run slower  $\Delta t' = \gamma \Delta t$
- Length contraction
  - moving objects appear contracted  $L' = \frac{L}{\gamma}$

- REALLY??? Can we check this experimentally???

Application:

## Cosmic Ray Muons

- **Cosmic ray muons:**
  - Cosmic rays are energetic particles (mainly protons) coming from somewhere in the Universe
  - When they hit the atmosphere they will produce showers of particles
  - $\mu$  are of particular interest because they are very penetrating and have a long lifetime ( $2.2 \mu\text{s}$ )
- **Question: Can muons produced in the upper atmosphere reach the ground?**
  - **Input:**
    - Muon's velocity = 99.99% of velocity of light  $c$
    - Atmosphere  $\sim 20 \text{ Km}$  thick

Application:

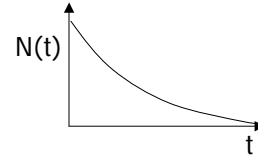
## Cosmic Ray Muons (2)

- **Inputs:**
  - $v_{\mu} = 99.99\%$  of velocity of light  $c$ , atmosphere  $\sim 20 \text{ Km}$
- **Non relativistic approach:**
  - $\Delta l = 0.9999 c \Delta t = 0.6 \text{ Km} < 20 \text{ Km}$ : **NO, they cannot reach the ground**
- **Relativistic approach**
  - $\gamma = 1/\sqrt{1-v^2/c^2} \sim 71$
  - **Approach 1: our perspective**
    - $\tau_{\mu} = 2.2 \mu\text{s}$  in muon's reference frame
    - In our reference frame:  $\tau' = \tau/\gamma = 71 \times 2.2 \mu\text{s} = 156 \mu\text{s}$
    - Now muon can travel:  $\Delta l = 42 \text{ Km}$ : **OK!**
  - **Approach 2: muons' perspective**
    - The  $\Delta l' = 20 \text{ Km}$  of atmosphere appear contracted to a relativistic  $\mu$
    - $\Delta l = \Delta l'/\gamma = 20\text{Km}/71 \sim 0.3 \text{ Km}$  that can be traveled with  $\tau=2.2 \mu\text{s}$ : **OK!**

Relativity: same physics in all reference frames!

## More on Cosmic Ray Muons

- The number of cosmic muons detected at sea level and on the top of Mount Everest are different. By how much?
  - Hypotheses:
    - Muons are produced in the upper atmosphere:  $\sim 20$  Km
    - $\beta = 0.9999 \rightarrow \gamma = 1/\sqrt{1-v^2/c^2} \sim 71$
    - Mount Everest  $\sim 8$  Km
    - Muons decay exponentially  $N(t) = N_0 \exp(-t/\tau)$
  - Choose 1 RF and stay with it
    - $\tau'_\mu = 156 \mu\text{s}$  in our R.F.
    - At sea level:
      - $L=20\text{Km} \rightarrow T=66 \mu\text{s} \rightarrow N_{\text{sea}} = N_0 \exp(-66/156) = 0.65 N_0$
    - On Mount Everest:
      - $L=12\text{Km} \rightarrow T=40 \mu\text{s} \rightarrow N_{\text{Everest}} = N_0 \exp(-40/156) = 0.77 N_0$
- $\rightarrow$  At sea level expect  $\sim 15\%$  less cosmic  $\mu$  than on Mount Everest: OK!



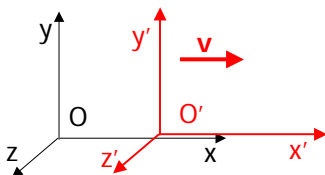
## How do lengths perpendicular to $v$ transform?

- Thought experiment
    - Train moving towards a tunnel with velocity  $v=0.9c$
    - Height of train in train's RF:  $h_{\text{train}} = 3.5$  m
    - Height of tunnel in tunnel's RF:  $h'_{\text{tunnel}} = 4.0$  m
  - If we have Lorentz contractions:  $L' = L/\gamma$ 
    - $\gamma = 1/\sqrt{1-0.9^2} = 2.29$
    - In tunnel's reference frame: the train moves with  $\beta=0.9$ 
      - $\rightarrow h'_{\text{train}} = h_{\text{train}}/\gamma = 3.5/2.29 = 1.5\text{m} \rightarrow$  no problem: it will fit!
    - In train's reference frame: tunnel moves with velocity  $\beta=0.9$ 
      - $\rightarrow h_{\text{tunnel}} = h'_{\text{tunnel}}/\gamma = 4/2.29 = 1.7\text{m} < h_{\text{train}} \rightarrow$  they will smash!
- $\rightarrow$  Different observers come to different conclusions  
 $\rightarrow$  against relativity principle!  $\rightarrow$  Lorentz contraction cannot happen



## Lorentz transformation

- "Time dilation" and "Length contraction" are consequences of the so called "Lorentz transformation"
- Consider 2 inertial reference frames:  $O$  and  $O'$ 
  - $O'$  is moving w.r.t.  $O$  with velocity  $\mathbf{v}$  //  $x$  axis where
    - $(x,y,z,t)$  the coordinate in the  $O$  reference frame
    - $(x',y',z',t')$  the coordinate in the  $O'$  reference frame



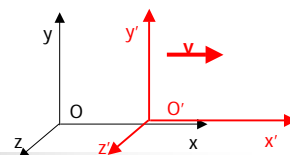
- Lorentz transformation:
  - Linear transformation that relates the coordinate in the 2 R.F.
    - Why linear? Because reference frames are inertial

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## Lorentz transformation (2)



- The most general form for a linear transformation:

- $z$  and  $y$  do not change because  $\mathbf{v}$  //  $x$   
 → ignore them in the following

$$\begin{cases} x' = Ax + Bt & (1) \\ y' = y \\ z' = z \\ t' = Cx + Dt & (2) \end{cases}$$

- Goal: calculate coefficients  $A, B, C, D$

- First requirement:

- $O$  and  $O'$  overlap at  $t=0$ : At  $t=t'=0, x=x'=0$

- For  $O$ , the origin of  $O'$  moves away with velocity  $\mathbf{v}$  →

Substitute in (1)  $\Rightarrow 0 = Avt + Bt \Rightarrow \boxed{B = -vA} \Rightarrow \begin{cases} x' = A(x - vt) & (3) \\ t' = Cx + Dt & (4) \end{cases}$

- For  $O'$ , the origin of  $O$  moves away with velocity  $-\mathbf{v}$  →

Substitute in (3):  $x' = A(x - vt) = -Avt$ . From (4):  $x' = -vt' = -v(Cx + Dt) - vDt$

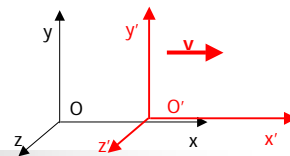
$$\Rightarrow \boxed{D = A} \Rightarrow \begin{cases} x' = A(x - vt) & (3) \\ t' = Cx + At & (5) \end{cases}$$

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## Lorentz transformation (3)



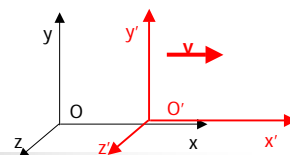
- **Second requirement:**

- Send a light pulse along the x direction at  $t=0$
- After a time  $t$  the coordinates of the light pulse are  $x=ct$  and  $x'=ct'$ . Substitute in (3) and use (5):

$$\begin{cases} ct' = x' = A(x - vt) = A(ct - vt) \\ ct' = c(Cx + At) = c(Cct + At) \end{cases} \Rightarrow c(Cct + At) = A(ct - vt) \Rightarrow C = -A \frac{v}{c^2}$$

$$\Rightarrow \begin{cases} x' = A(x - vt) & (3) \\ t' = A\left(t - \frac{v}{c^2}x\right) & (6) \end{cases}$$

## Lorentz transformation (4)



- **Third requirement:**

- Send a light pulse along the y direction at  $t=0$
- After a time  $t$  the coordinates of the light pulse are  $(x=0; y=ct)$  in  $O$ ; in  $O'$  the total displacement is:  $x'^2 + y'^2 = (ct')^2$ . Substitute (3) and (6):

$$x'^2 + y'^2 = (ct')^2$$

$$A^2(x - vt)^2 + y^2 = c^2 A^2 \left( t - \frac{v}{c^2} x \right)^2$$

$$\text{Since } x=0 \text{ and } y=ct \Rightarrow A^2(vt)^2 + (ct)^2 = c^2 A^2 t^2$$

$$\Rightarrow A = \frac{1}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \equiv \gamma$$

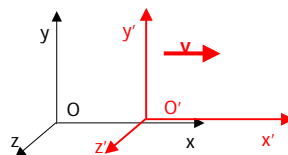
$$\Rightarrow \begin{cases} x' = \gamma(x - vt) \\ t' = \gamma\left(t - \frac{v}{c^2}x\right) \end{cases}$$

## Lorentz transformation: summary

Summarizing: when  $O'$  moves wrt  $O$  with velocity  $+v//x$  axis

- To go from  $O$  (at rest) to  $O'$  (in motion):

$$\begin{cases} x' = \gamma(x - vt) \\ t' = \gamma\left(t - \frac{v}{c^2}x\right) \end{cases}$$

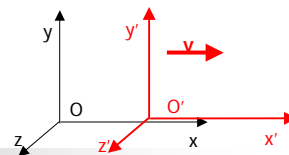


- To go from  $O'$  (in motion) to  $O$  (at rest), just change the sign of the velocity:

$$\begin{cases} x = \gamma(x' + vt') \\ t = \gamma\left(t' + \frac{v}{c^2}x'\right) \end{cases}$$

- The other coordinates ( $y$  and  $z$ ) are not affected

## Transformation of velocity



- Consequence of Lorentz transformations
- Observer in motion  $O'$  shoots a bullet with velocity  $u'_x // +x$  axis
- What is the velocity of the bullet  $u_x$  measured by  $O$ ?

$$\begin{aligned} u_x &= \frac{dx}{dt} = \frac{d(\gamma(x' + vt'))}{d(\gamma(t' + \frac{v}{c^2}x'))} = \frac{dx' + v dt'}{dt' + \frac{v}{c^2} dx'} \\ &= \frac{dx' / dt' + v}{1 + \frac{v}{c^2} dx' / dt'} = \frac{u'_x + v}{1 + \frac{vu'_x}{c^2}} \end{aligned}$$

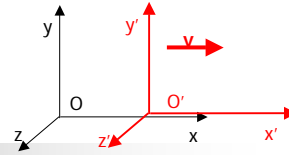
- Conclusion:

$$u_x = \frac{u'_x + v}{1 + \frac{vu'_x}{c^2}}$$

and

$$u'_x = \frac{u_x - v}{1 - \frac{vu_x}{c^2}}$$

## Velocity not // to $v$



- How do we sum velocity not // to the relative motion of the 2 R.F.?
- Observer in motion  $O'$  shoots a bullet with velocity  $u'_y$  perpendicular to  $v$
- What is the velocity of the bullet  $u_x$  measured by  $O$ ?

$$u_y = \frac{dy}{dt} = \frac{dy'}{d\left(\gamma\left(t' + \frac{v}{c^2}x'\right)\right)} = \frac{dy'}{\gamma\left(dt' + \frac{v}{c^2}dx'\right)}$$

$$= \frac{dy'/dt'}{\gamma\left(dt' + \frac{v}{c^2}dx'\right)/dt'} = \frac{u'_y}{\gamma\left(1 + \frac{vu'_x}{c^2}\right)}$$

- Conclusion:

$$u_y = \frac{u'_y}{\gamma\left(1 + \frac{vu'_x}{c^2}\right)} \quad \text{and} \quad u'_y = \frac{u_y}{\gamma\left(1 - \frac{vu_x}{c^2}\right)}$$

## Summary and outlook

- Today:
  - Principle of Special Relativity and its amazing consequences
    - Length contraction and Time dilation
    - Lorentz transformations
    - Velocity transformation ( $v$  always  $< c$ )
- Next time:
  - More on Relativity:
    - How to transform electric fields and forces
    - Prove that  $E$  and  $B$  are intimately connected