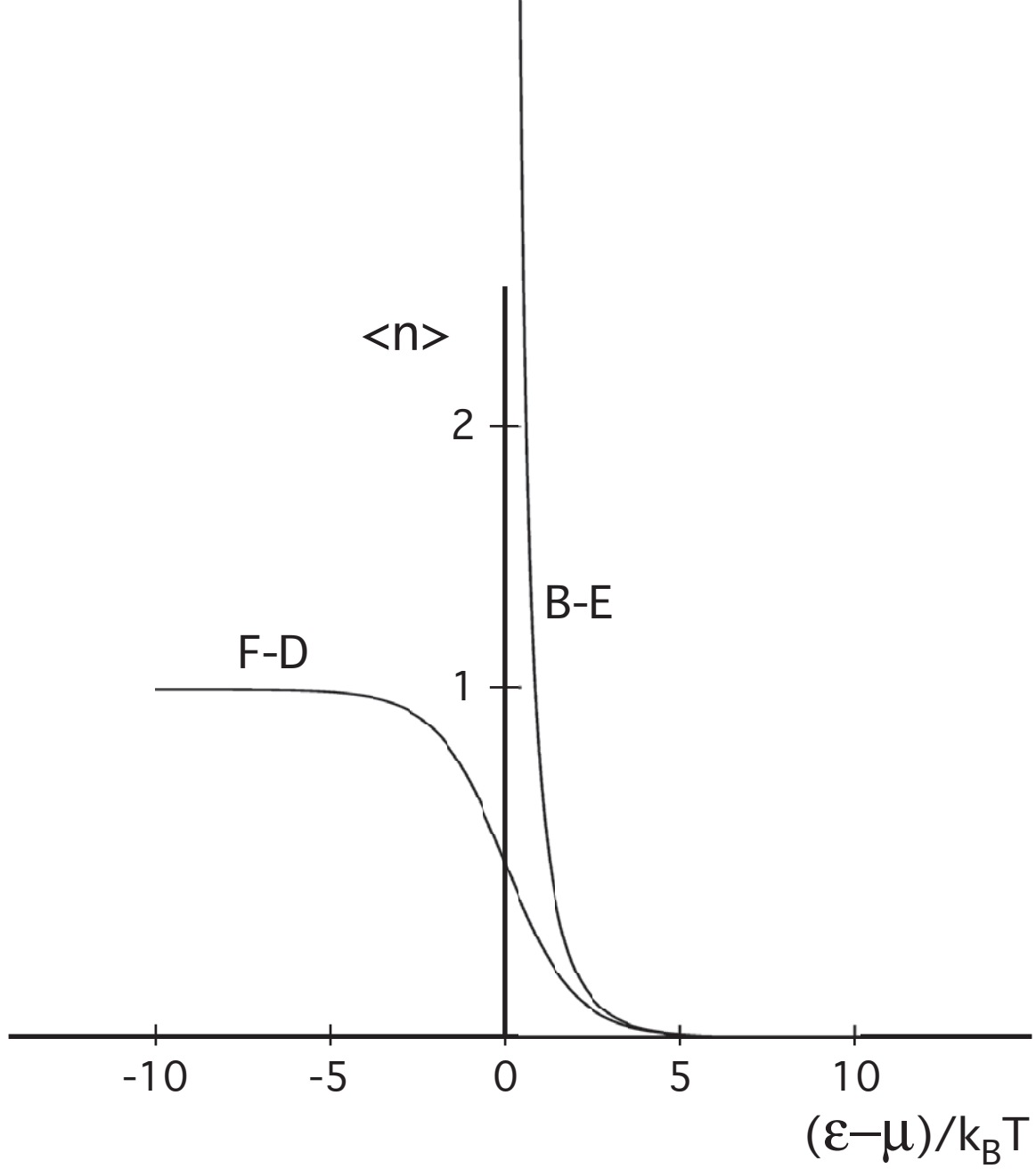


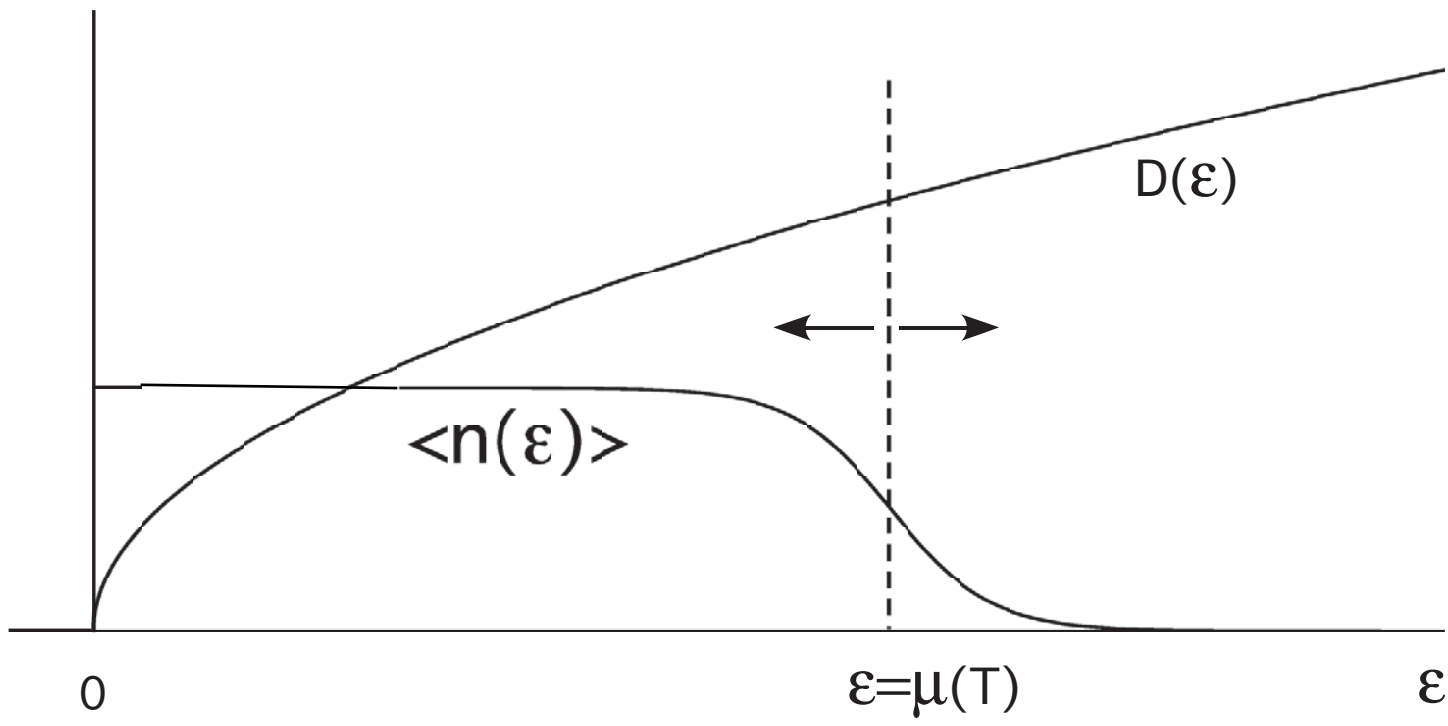
$$N = \int_0^{\infty} \langle n(\epsilon, \mu(T), T) \rangle D(\epsilon) d\epsilon$$

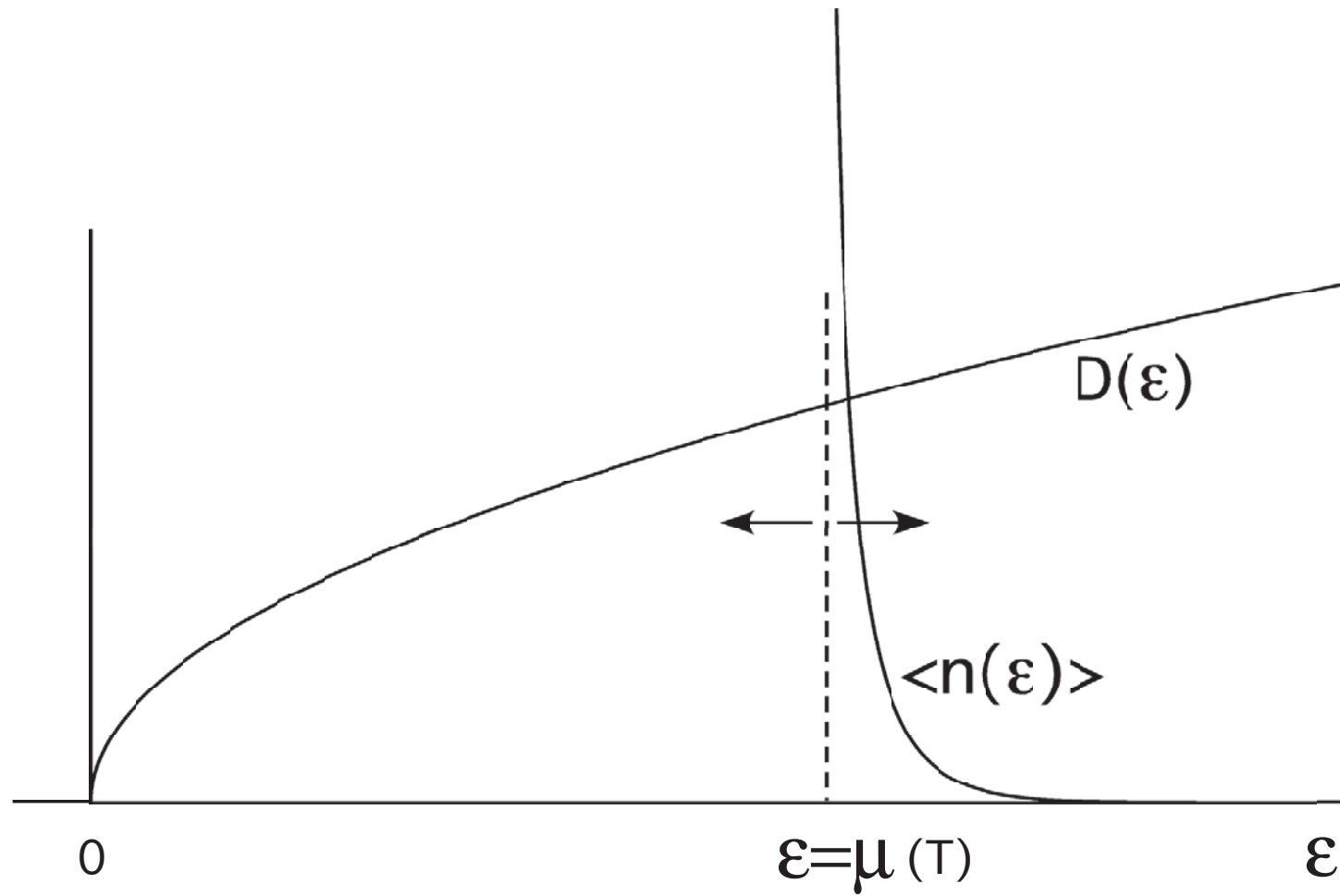
This expression implicitly determines  $\mu = \mu(T)$ .

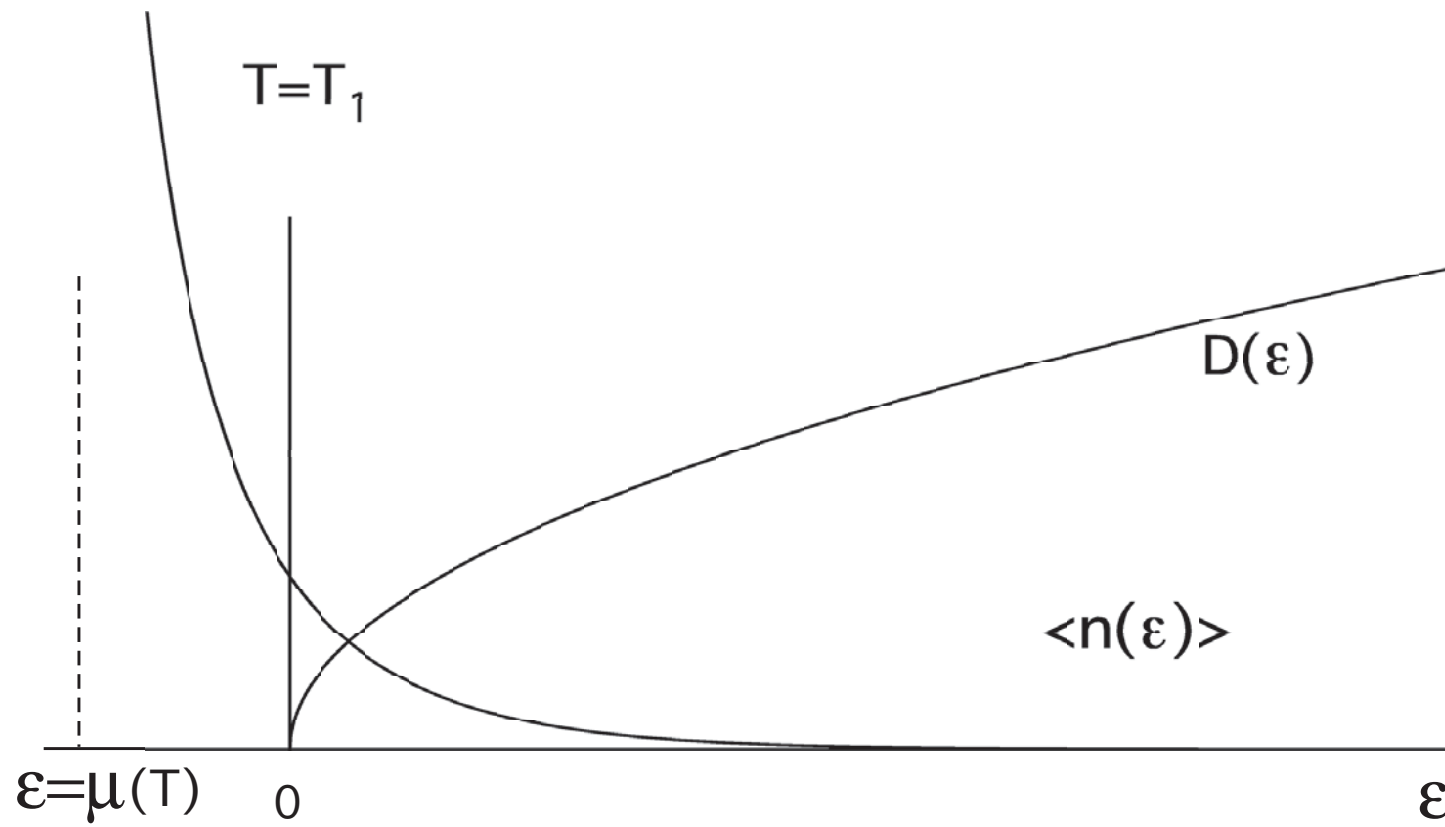
$$U = \int_0^{\infty} \langle n(\epsilon, \mu(T), T) \rangle \epsilon D(\epsilon) d\epsilon$$

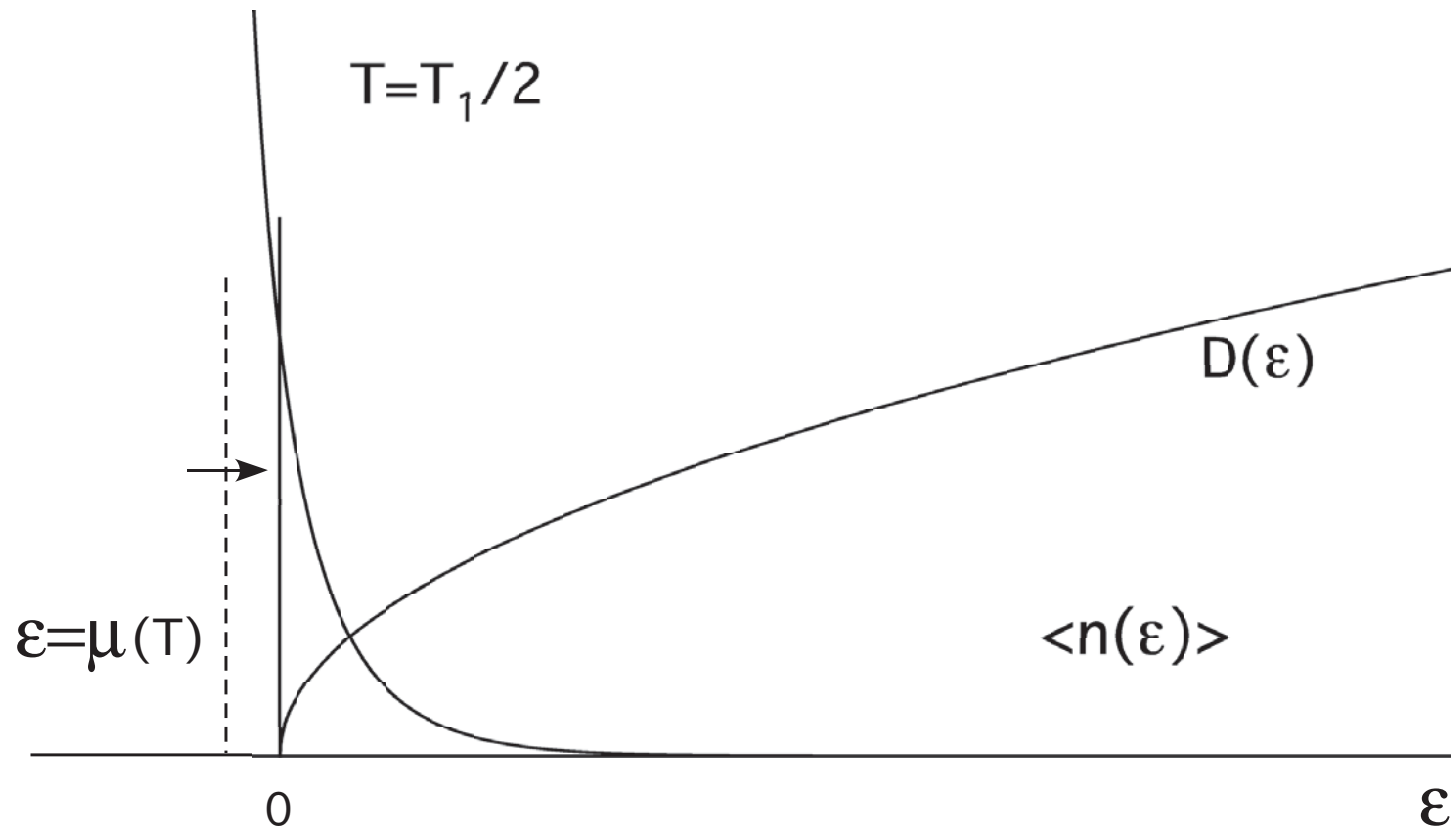
To determine  $C_V$  from this expression one must take into account the temperature dependence of  $\mu$  in addition to the explicit dependence of  $\langle n \rangle$  on  $T$ .

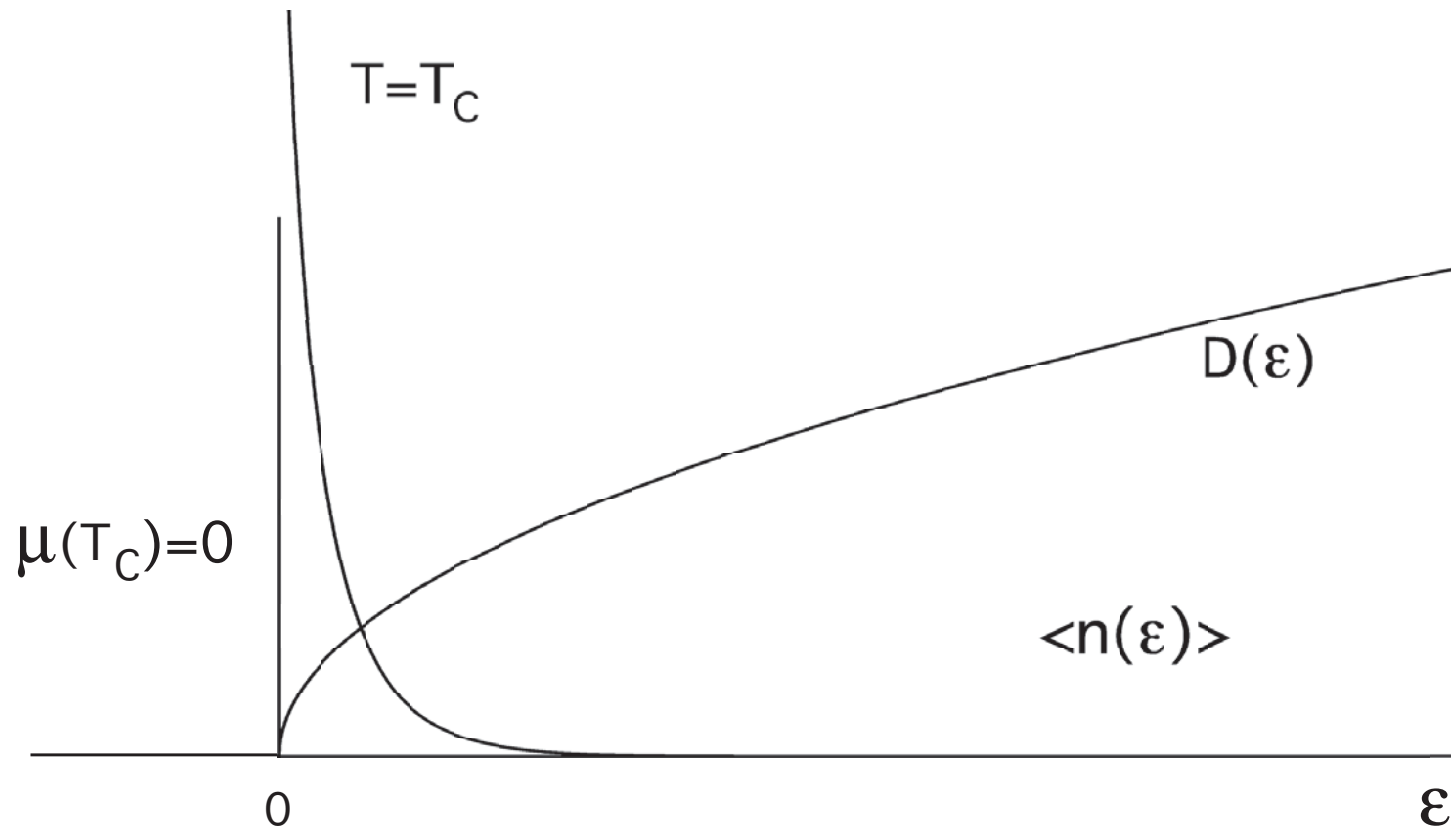












$$\begin{aligned}
N &= \int_0^\infty \langle n \rangle D(\epsilon) d\epsilon \\
&= \int_0^\infty \frac{1}{e^{\epsilon/k_B T} - 1} \left[ \frac{(2S + 1)V}{4\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \sqrt{\epsilon} \right] d\epsilon \\
&= \frac{(2S + 1)V}{4\pi^2} \left( \frac{2m}{\hbar^2} \right)^{3/2} \int_0^\infty \frac{\sqrt{\epsilon}}{e^{\epsilon/k_B T} - 1} d\epsilon \\
&= \frac{(2S + 1)V}{4\pi^2} \left( \frac{2mk_B T}{\hbar^2} \right)^{3/2} \underbrace{\int_0^\infty \frac{\sqrt{x}}{e^x - 1} dx}_{(\sqrt{\pi}/2)\zeta(3/2)} \\
&= (2S + 1)\zeta(3/2) \left( \frac{V}{\lambda^3(T)} \right)
\end{aligned}$$



$$\zeta(3/2) = 2.612\dots \quad \text{and} \quad \lambda(T) = h/\sqrt{2\pi mk_B T}.$$

For a fixed number density  $n = N/V$  the critical temperature for  $S = 0$  Bosons is given by

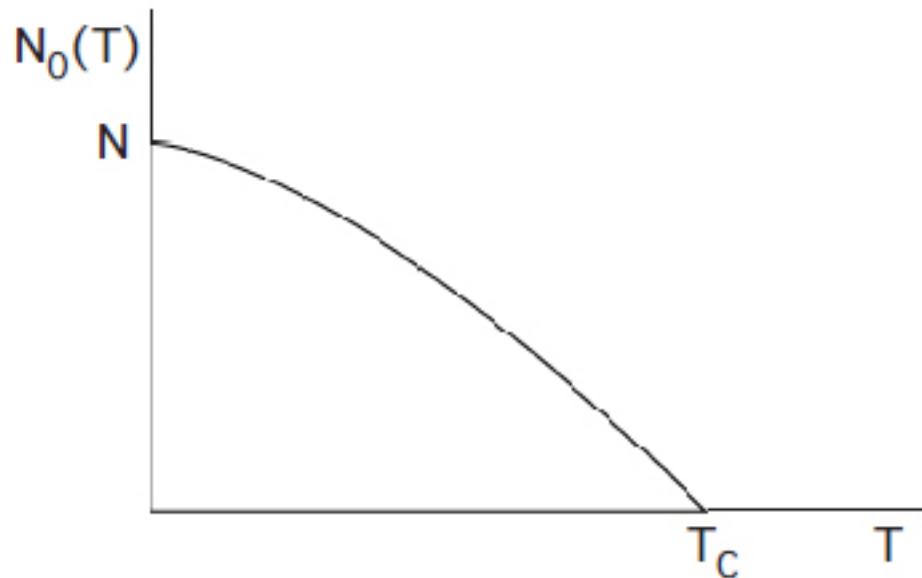
$$T_c = \frac{h^2}{2\pi mk_B} \left( \frac{n}{\zeta(3/2)} \right)^{2/3} \propto n^{2/3}$$

At a fixed temperature the critical number density is given by

$$n_c = \zeta(3/2) \lambda^{-3}(T) \propto T^{3/2}$$

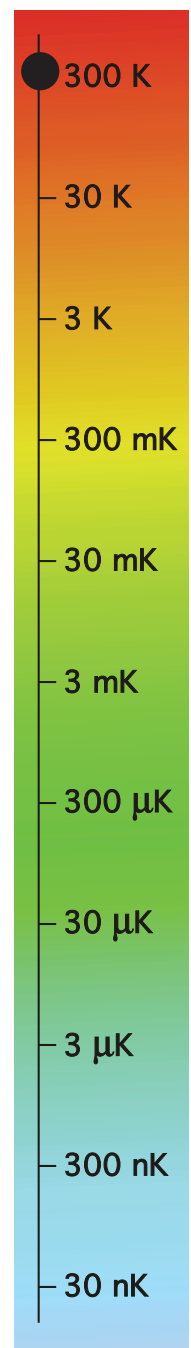
For a fixed number of atoms  $N$ , once  $T$  falls below  $T_c$  those atoms that can not be accommodated in states with finite  $\epsilon$  begin accumulating in the ground state where  $\epsilon = 0$ .

$$N_0 = N - \underbrace{\int_0^\infty \langle n \rangle D(\epsilon) d\epsilon}_{\propto T^{3/2}} = N(1 - (T/T_c)^{3/2})$$



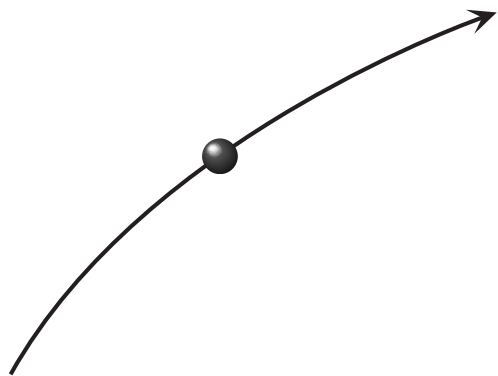


# BOSE-EINSTEIN CONDENSATION IS A QUANTUM MECHANICAL EFFECT



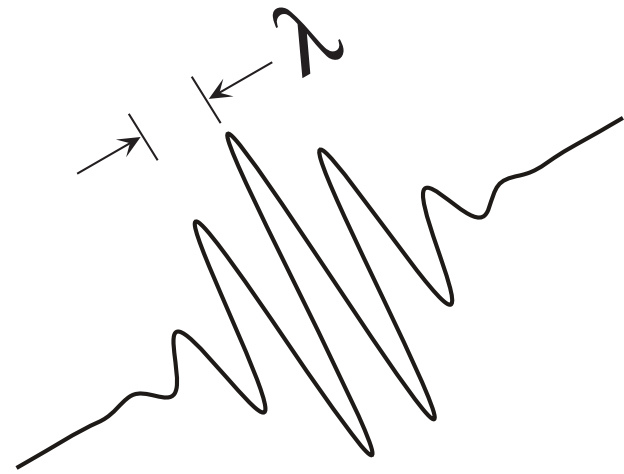


## CLASSICAL MODEL

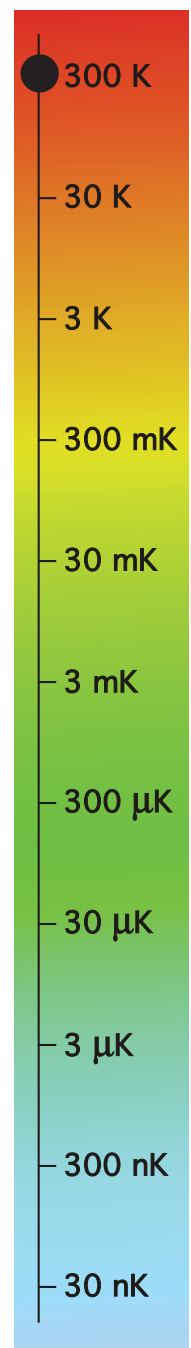


POINT-LIKE  
PARTICLES  
FOLLOWING  
TRAJECTORIES

## QUANTUM REALITY



WAVES  
PROPAGATING  
THROUGH  
SPACE



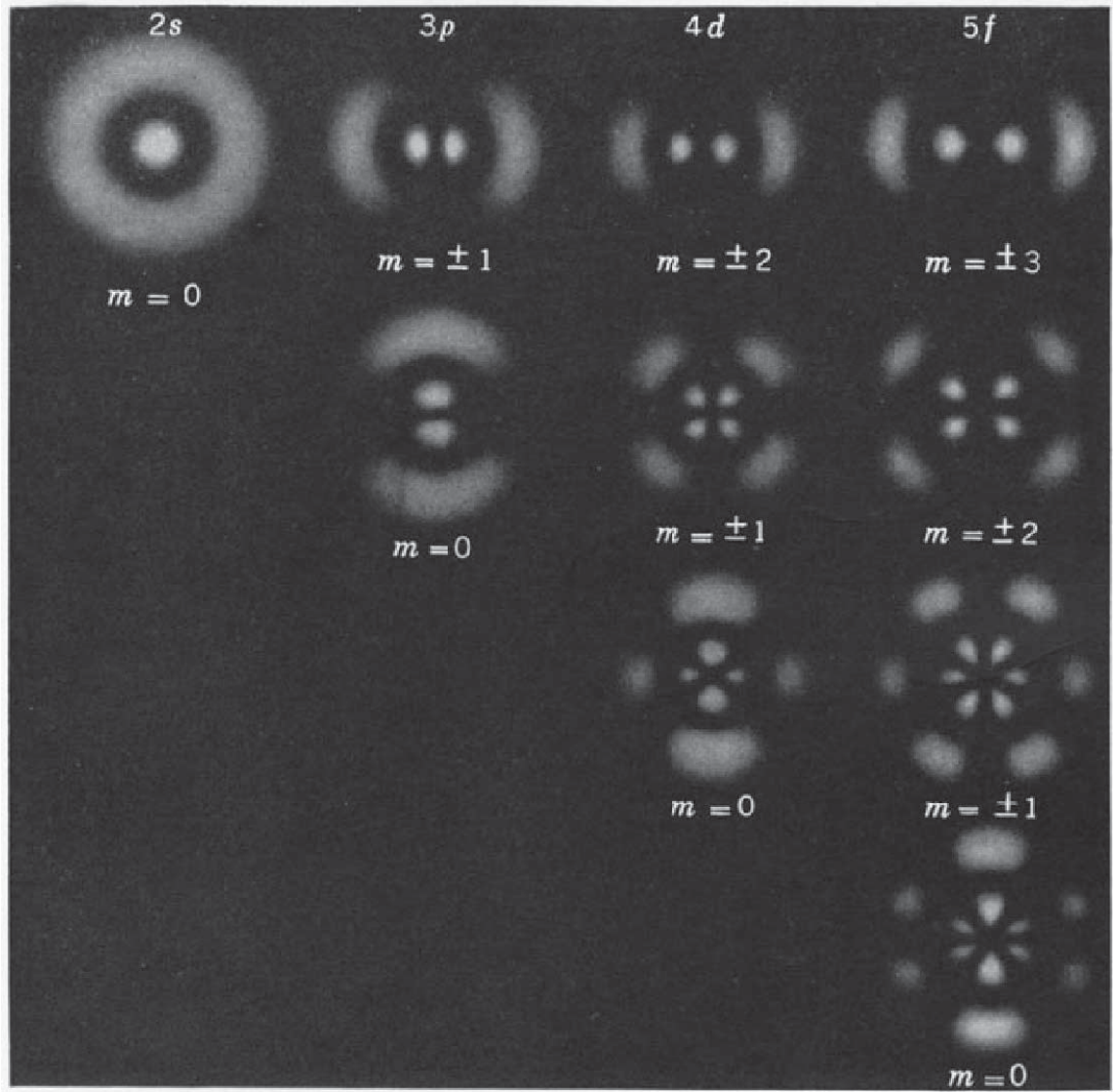
1900 1920 1940 1960 1980 2000

$$\lambda \propto \frac{1}{m \times V}$$

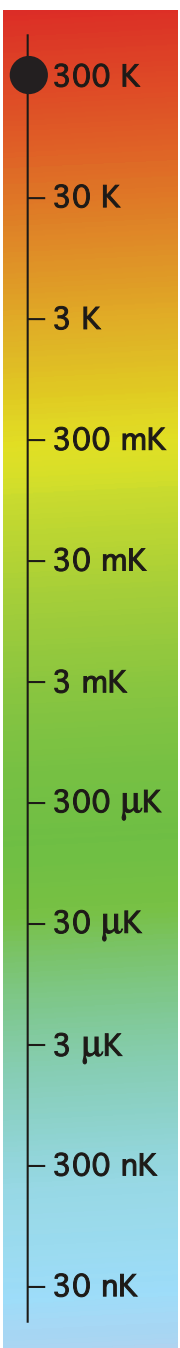
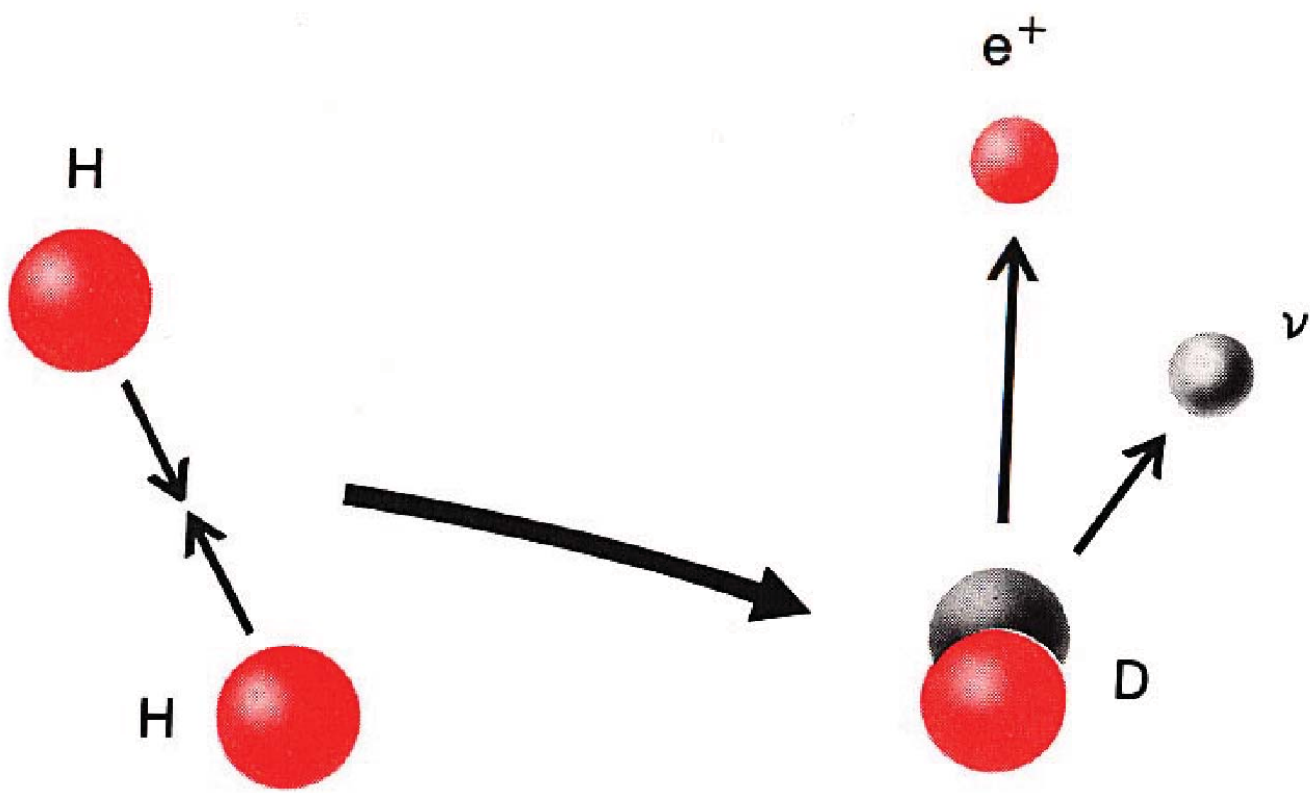
FOR ATOMS MOVING AT THERMAL VELOCITY  
AT ROOM TEMPERATURE (300K),  
 $\lambda <$  THEIR PHYSICAL SIZE.

FOR THE ELECTRONS MOVING AROUND THE NUCLEI  
IN THOSE ATOMS,  $\lambda \approx 1$  ANGSTROM.

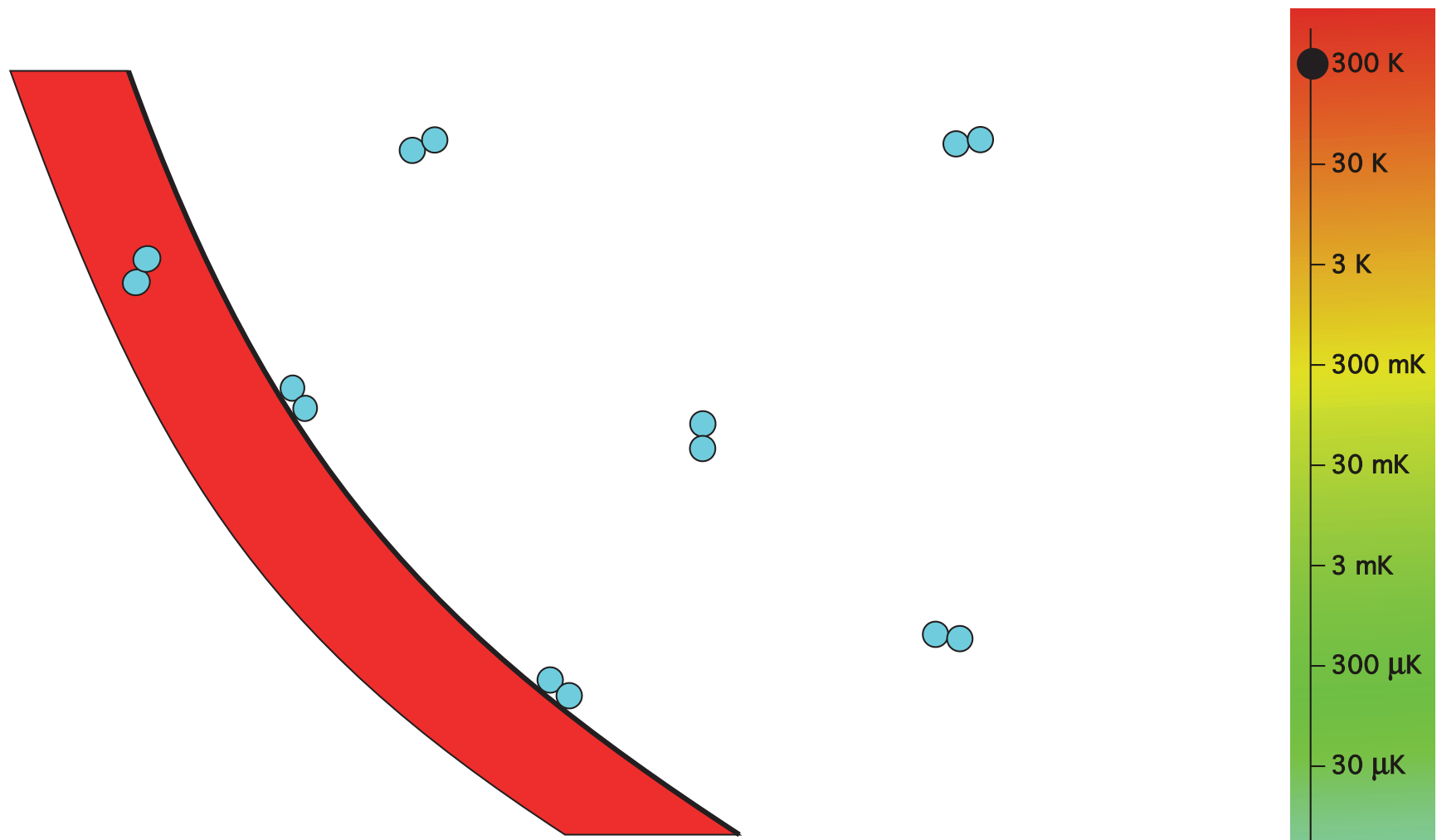




THE WAVE NATURE OF THE ELECTRONS  
 STABILIZES THEM AGAINST LOSING ENERGY AND  
 FALLING INTO THE NUCLEUS.



THE WAVE NATURE OF PROTONS ALLOWS THEM TO GET CLOSE ENOUGH DURING COLLISIONS IN THE SUN TO INITIATE FUSION.



QM ALLOWS MOLECULES TO HAVE A STATISTICAL CHANCE OF ADSORBING ON A SURFACE INSTEAD OF REMAINING IN THE BULK GAS



1900

1920

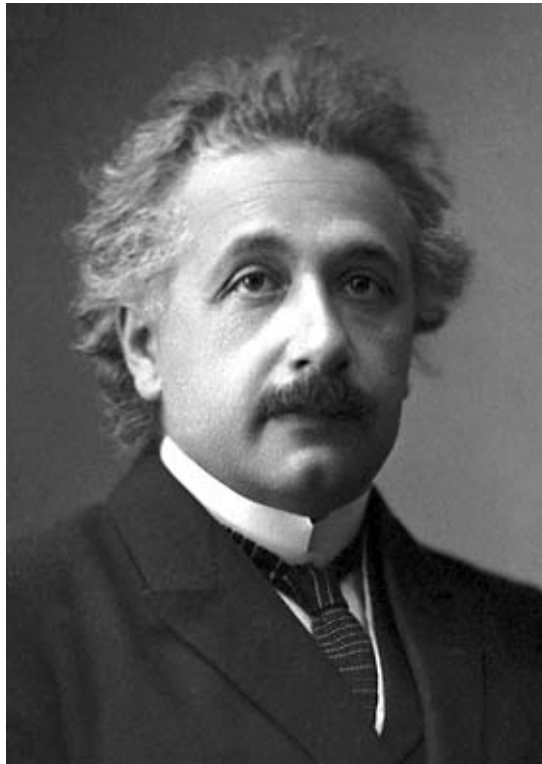
1940

1960

1980

2000

IN 1924 AND 1925 SATYENDRA BOSE AND ALBERT EINSTEIN INVESTIGATED THE INFLUENCE OF QM ON THE COLLECTIVE BEHAVIOR OF PARTICLES.



1921

"for his services to Theoretical Physics, and especially for his discovery of the law of the photoelectric effect"

300 K

30 K

3 K

300 mK

30 mK

3 mK

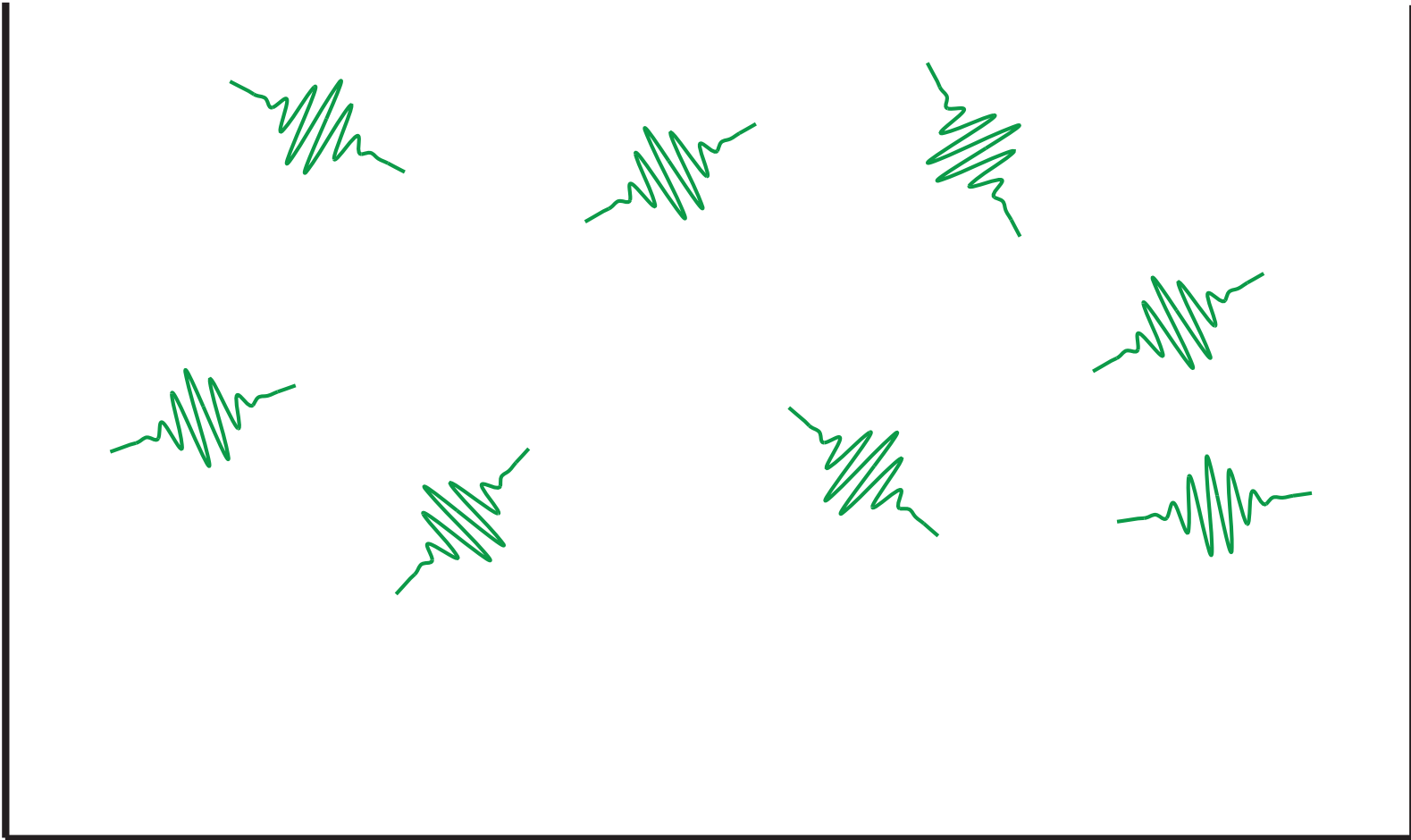
300  $\mu$ K

30  $\mu$ K

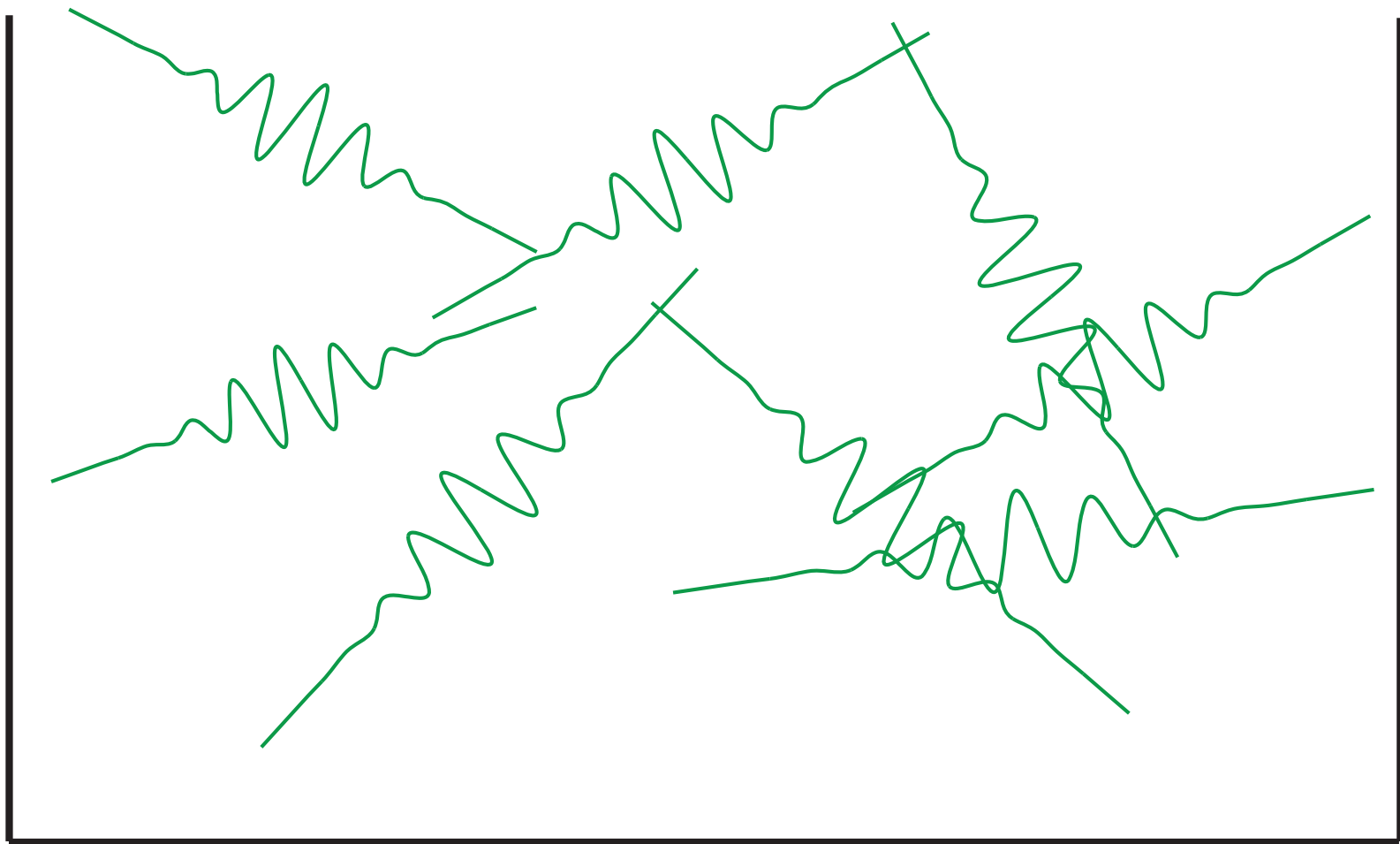
3  $\mu$ K

300 nK

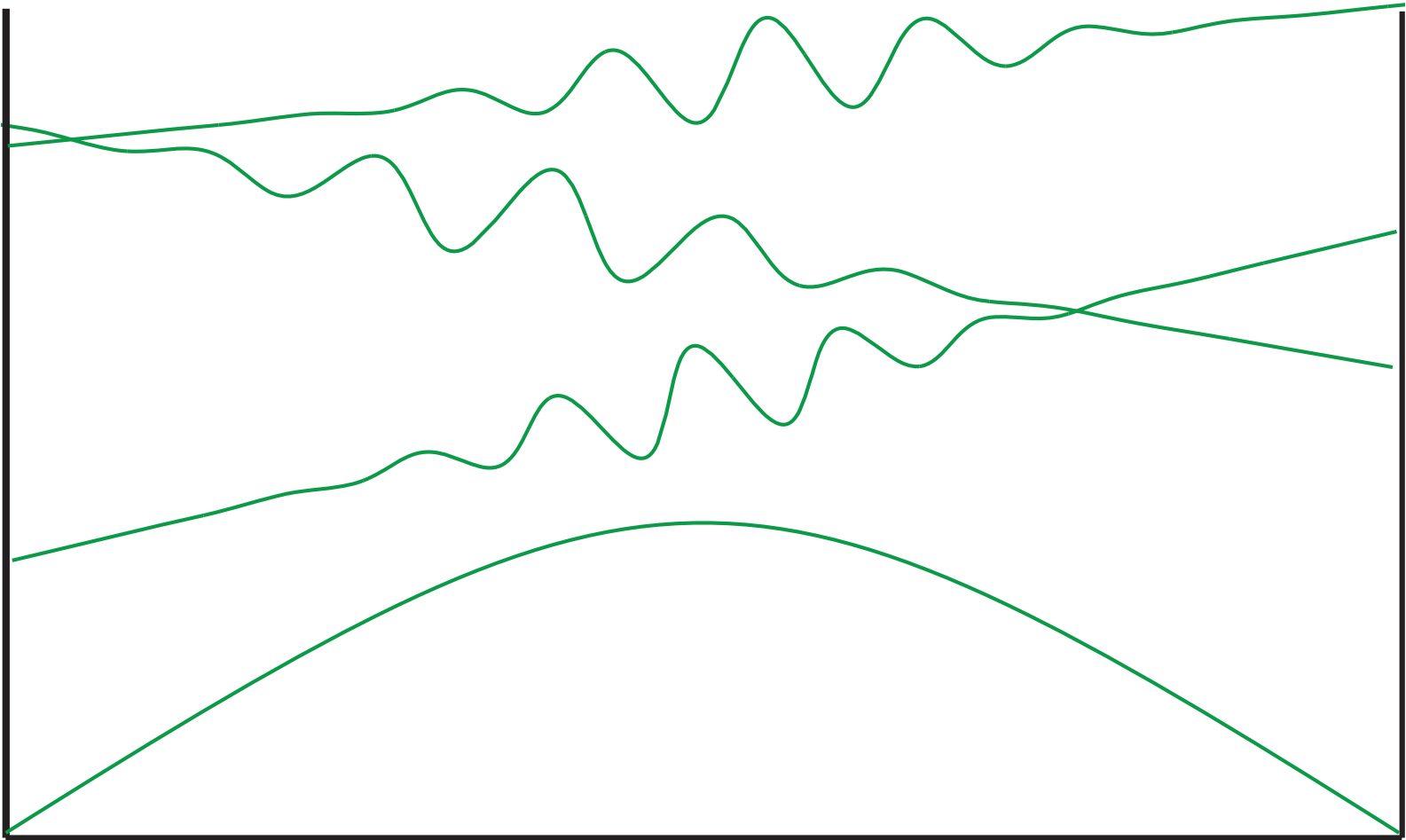
30 nK



THEY APPLIED THEIR THEORY TO THE SIMPLEST POSSIBLE CASE: A GAS OF NON-INTERACTING ATOMS



AS THE ATOMS GET COLDER THEIR VELOCITY DIMINISHES AND THEIR WAVELENGTH GROWS.



WHEN THE WAVELENGTH BECOMES COMPARABLE TO THE SEPARATION, A PHASE TRANSITION OCCURS. SOME OF THE ATOMS LOSE THEIR IDENTITY AND BECOME PART OF A SINGLE WAVE SPANNING THE CONTAINER.

1900

1920

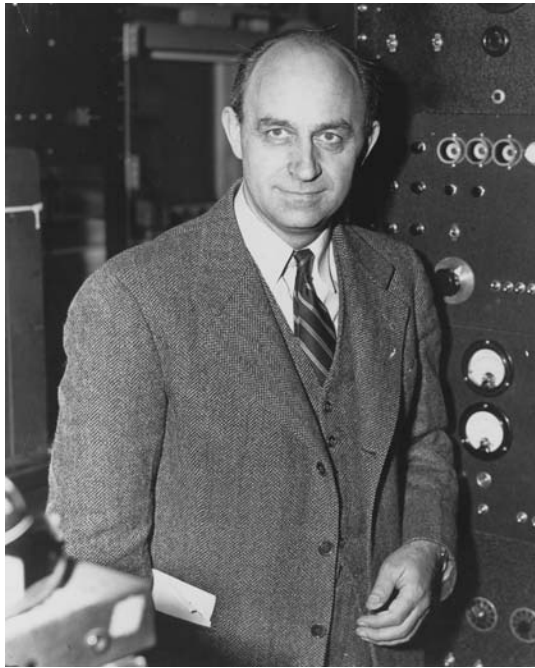
1940

1960

1980

2000

HALF THE ATOMS IN THE WORLD FOLLOW THE RULES OF BOSE AND EINSTEIN AND ARE CALLED "BOSONS". THE OTHER HALF FOLLOW RULES SET OUT BY ENRICO FERMI AND PAUL DIRAC AND ARE CALLED "FERMIONS".



Enrico Fermi 1938  
"for his demonstrations of the existence of new radioactive elements produced by neutron irradiation, and for his related discovery of nuclear reactions brought about by slow neutrons"



P.A.M. Dirac 1933 (with Erwin Schrodinger)  
"for the discovery of new productive forms of atomic theory"

300 K

30 K

3 K

300 mK

30 mK

3 mK

300  $\mu$ K30  $\mu$ K3  $\mu$ K

300 nK

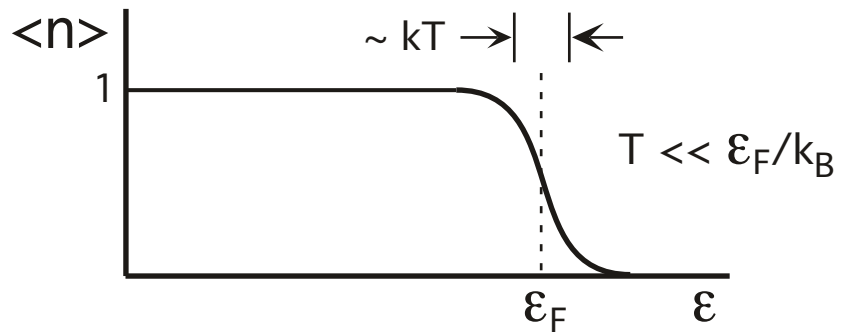
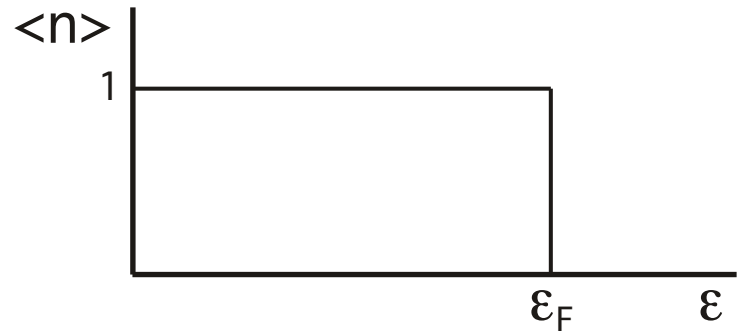
30 nK



$T = 0$

FINITE  $T$

FERMI

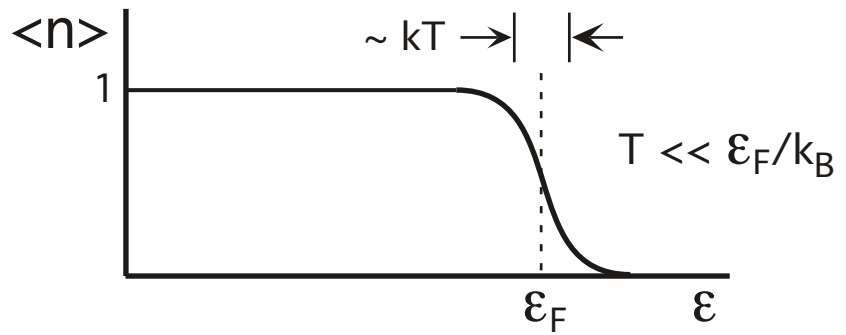
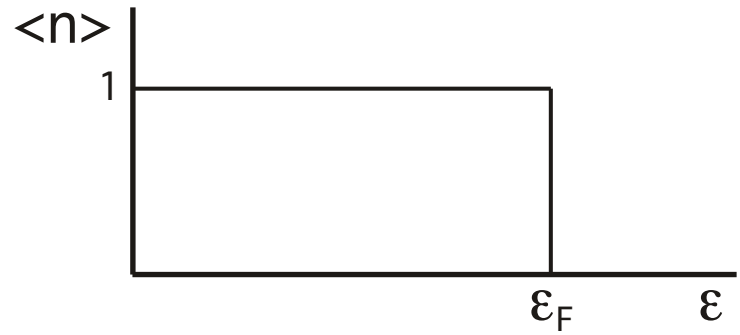




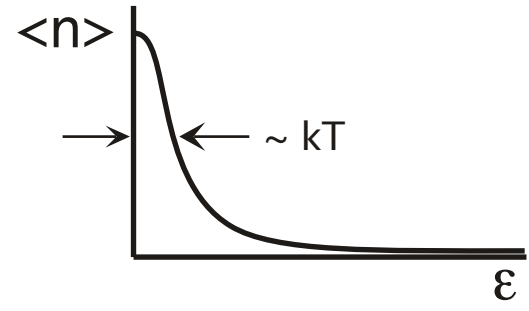
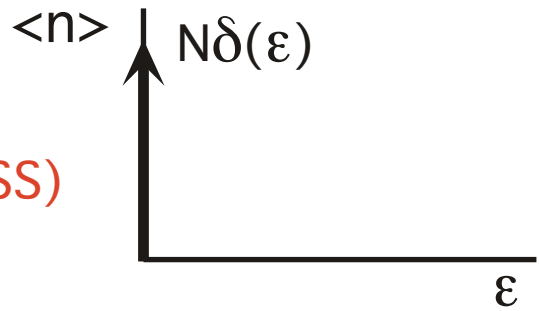
**T = 0**

**FINITE T**

**FERMI**



**BOSE**  
(GOOD GUESS)

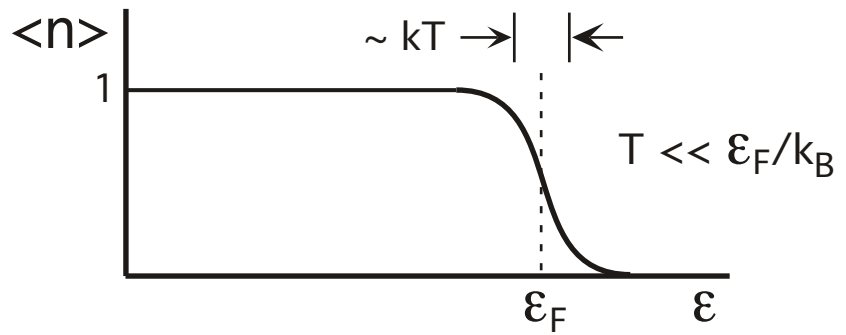
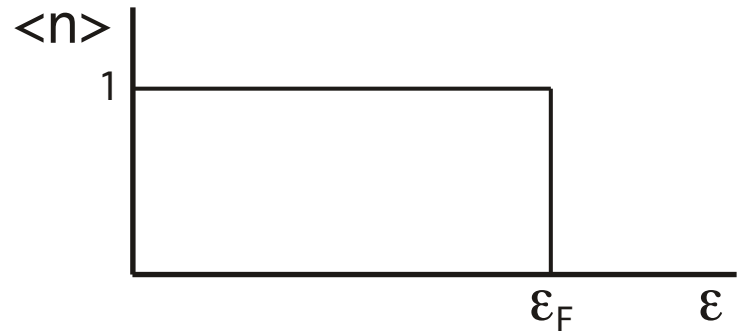




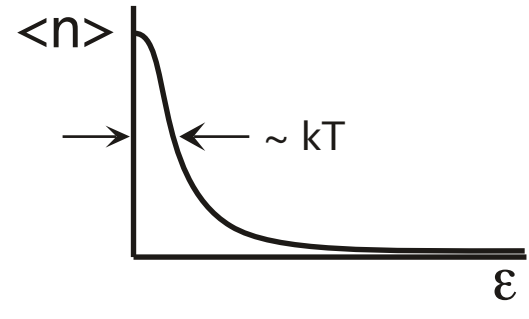
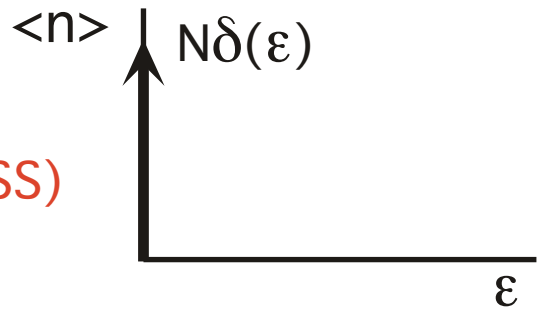
**T = 0**

**FINITE T**

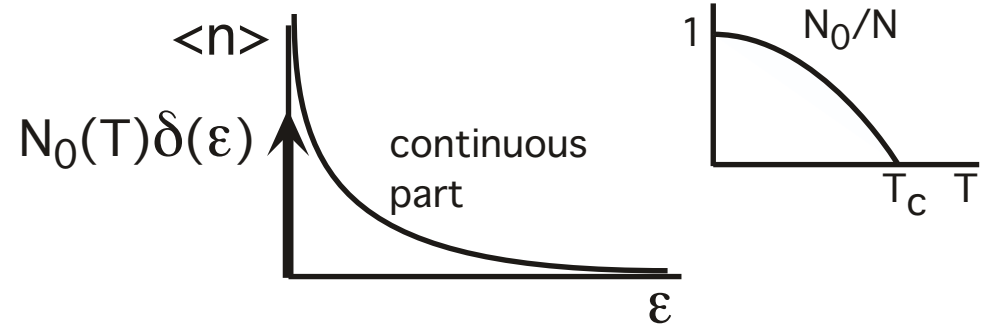
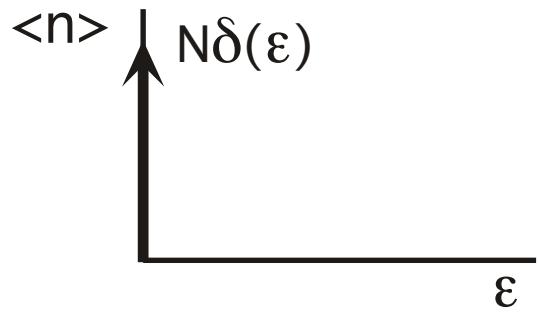
**FERMI**



**BOSE  
(GOOD GUESS)**



**BOSE  
(ACTUAL)**







O<sub>2</sub> LIQUEFIES AT 90K

O<sub>2</sub> FREEZES AT 50K

H<sub>2</sub> LIQUEFIES AT 20K

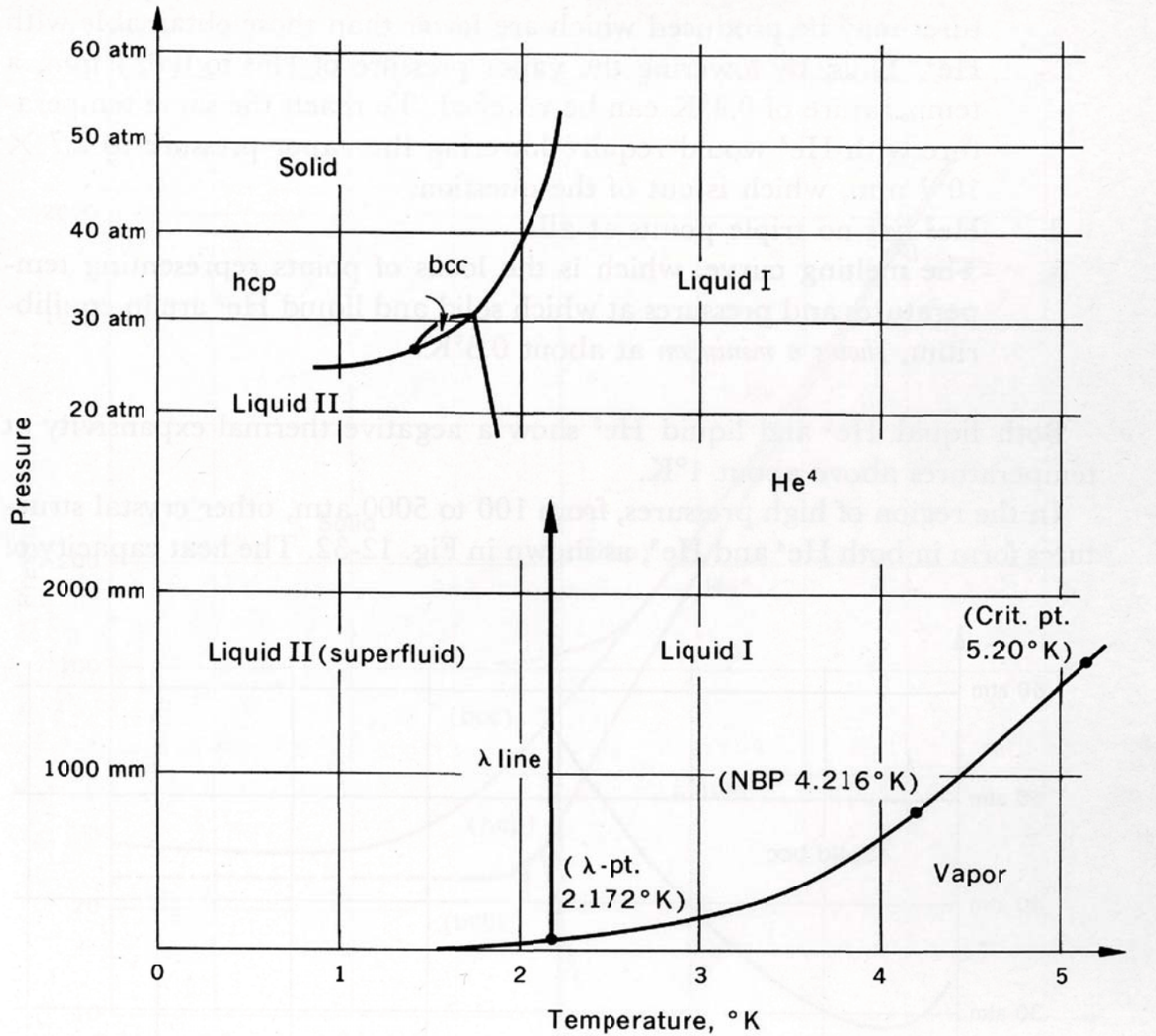
H<sub>2</sub> FREEZES AT 14K

He LIQUEFIES AT 4K



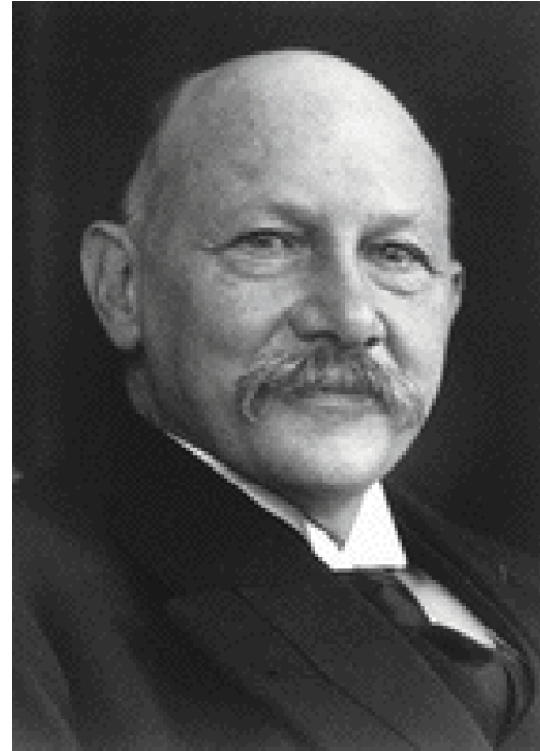
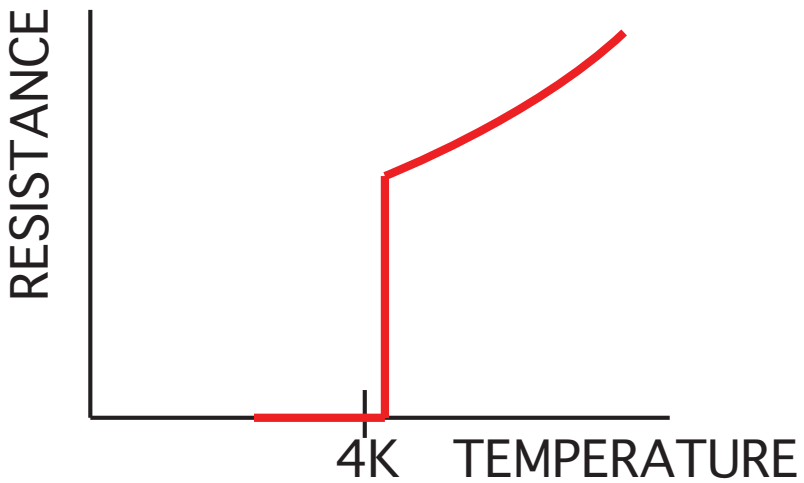
REAL ATOMS OR MOLECULES DO INTERACT WITH EACH OTHER AND UNDERGO LIQUEFICATION AND FREEZING DUE TO THESE INTERACTIONS.

# TEMPERATURES BELOW 4.2 K CAN BE ACHIEVED IN $^4\text{He}$ BY PUMPING ON THE VAPOR ABOVE THE LIQUID

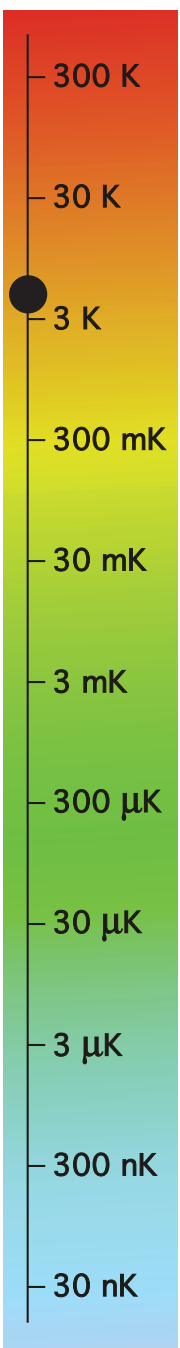




# SUPERCONDUCTIVITY WAS DISCOVERED BY KAMERLINGH ONNES IN 1911.



1913  
"for his investigations on the properties of matter at low temperatures which led, inter alia, to the production of liquid helium"



1900

1920

1940

1960

1980

2000



## The Nobel Prize in Physics 1972

The Nobel Prize in Physics 1972 was awarded jointly to [John Bardeen](#), [Leon Neil Cooper](#) and [John Robert Schrieffer](#) "for their jointly developed theory of superconductivity, usually called the *BCS-theory*".

300 K

30 K

3 K

300 mK

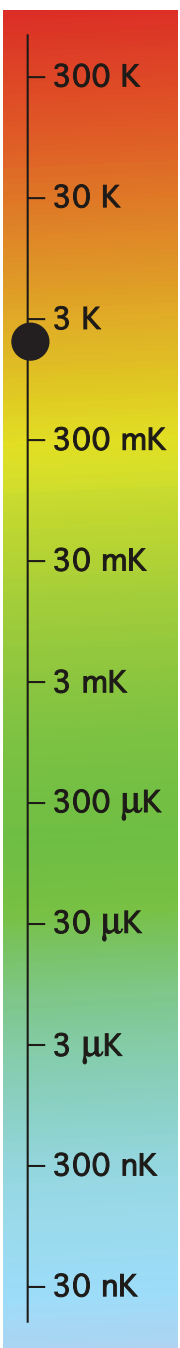
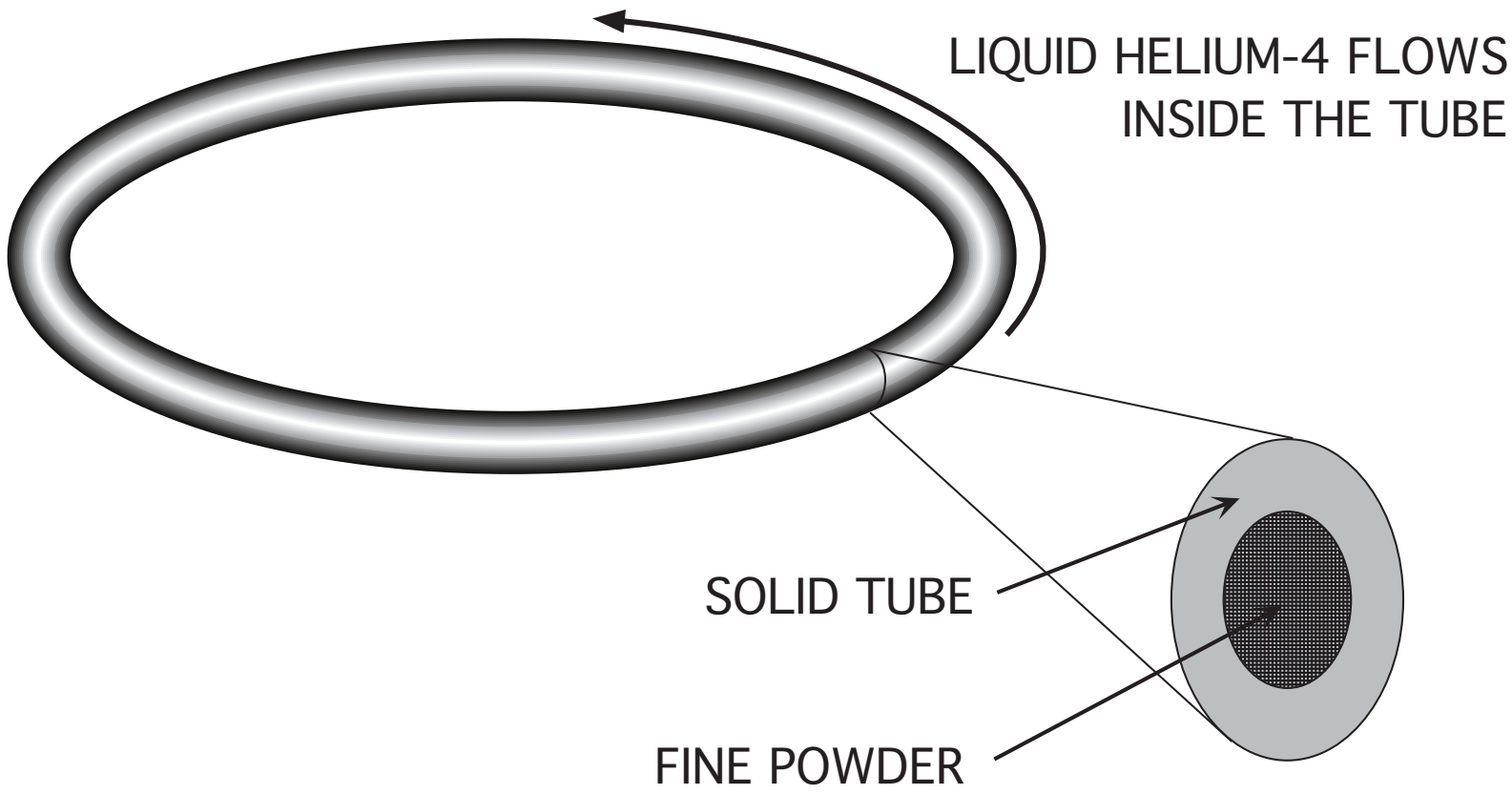
30 mK

3 mK

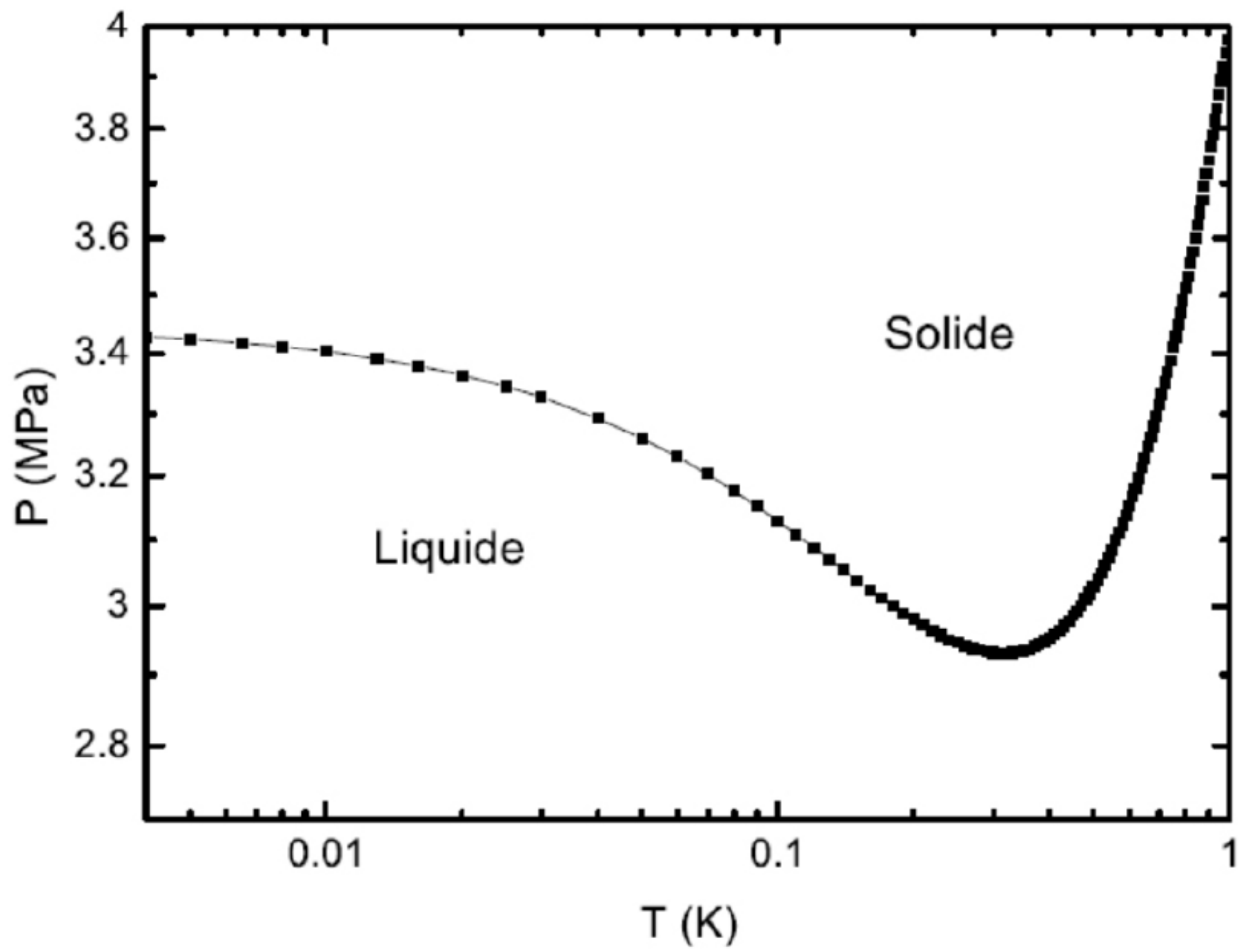
300  $\mu$ K30  $\mu$ K3  $\mu$ K

300 nK

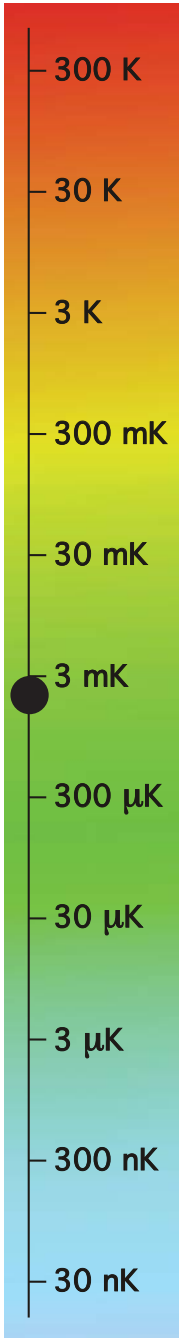
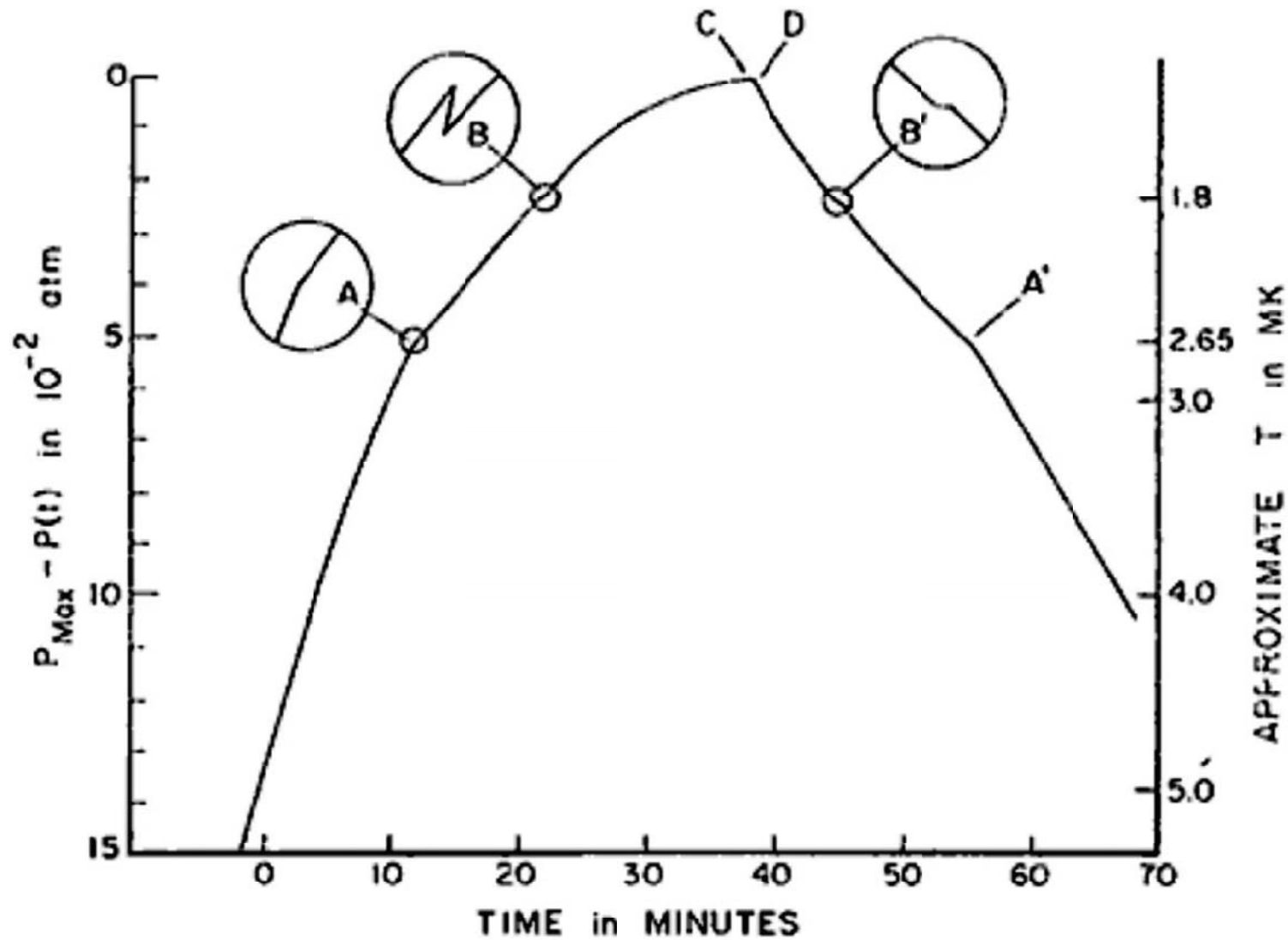
30 nK



SUPERFLUIDITY WAS DISCOVERED IN HELIUM-4 IN THE 1930s AT A TEMPERATURE OF 2 KELVIN.

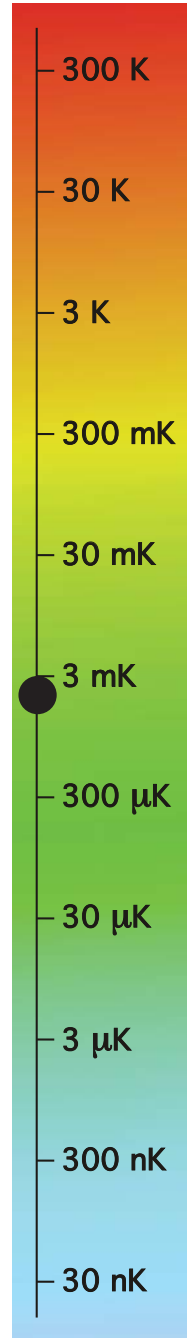
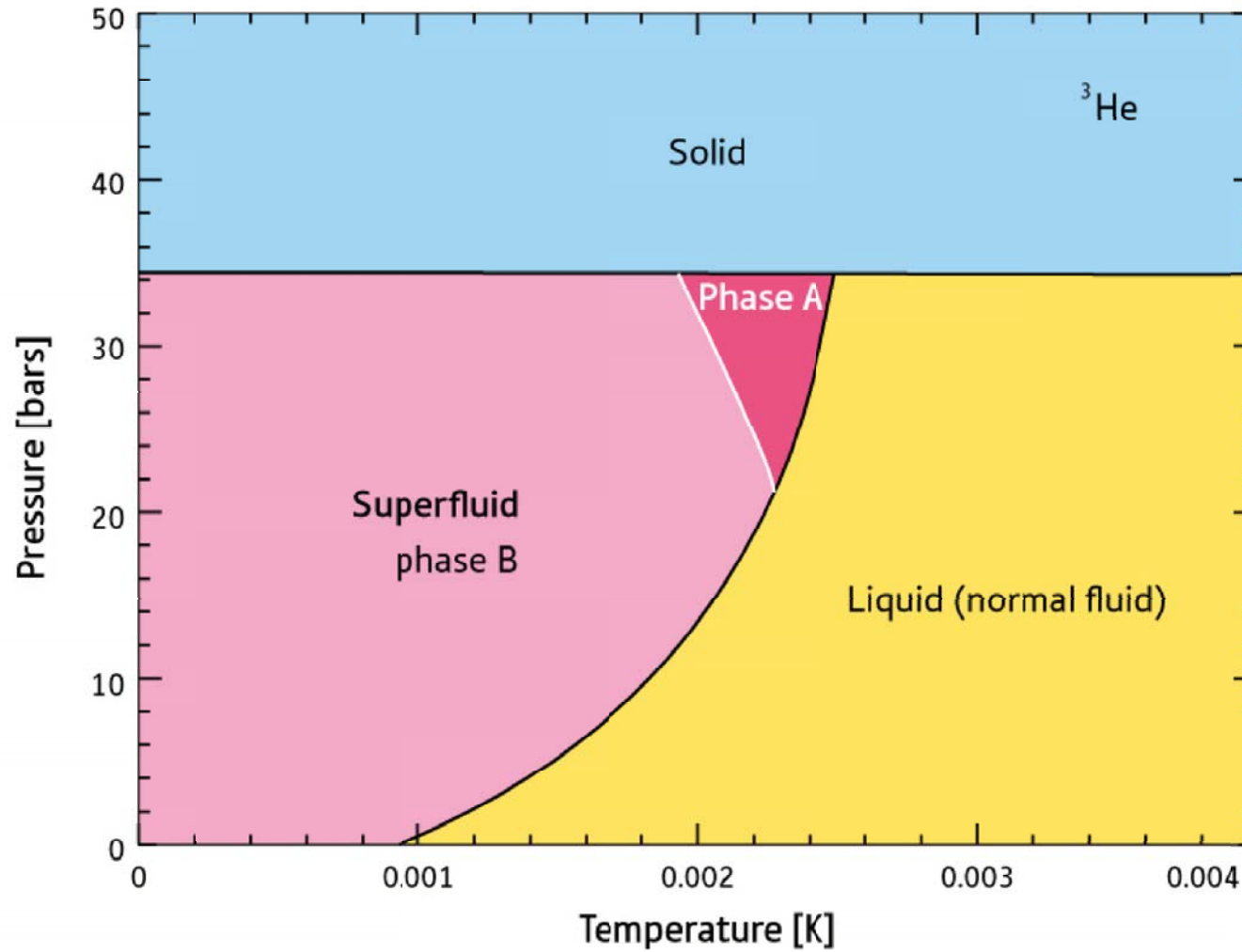


# DOUG OSHEROFF SEES STRANGE GLITCHES ON THE MELTING CURVE OF $^3\text{He}$



1900 1920 1940 1960 1980 2000

# WHICH TURNED OUT TO BE TWO DIFFERENT SUPERFLUID PHASES OF $^3\text{He}$

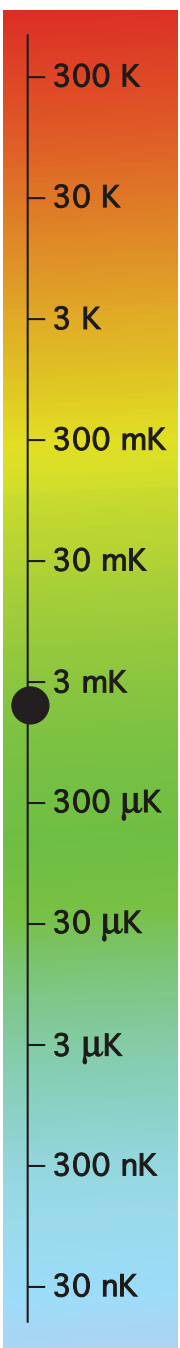






# The Nobel Prize in Physics 1996

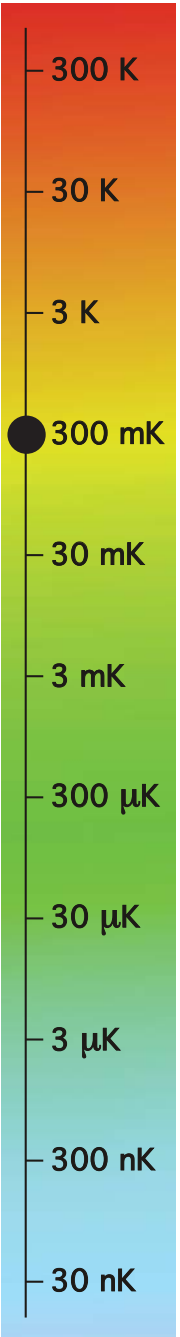
The Nobel Prize in Physics 1996 was awarded jointly to [David M. Lee](#), [Douglas D. Osheroff](#) and [Robert C. Richardson](#) "*for their discovery of superfluidity in helium-3*".

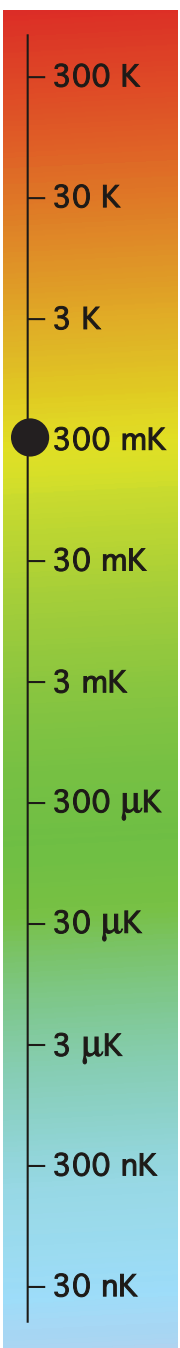
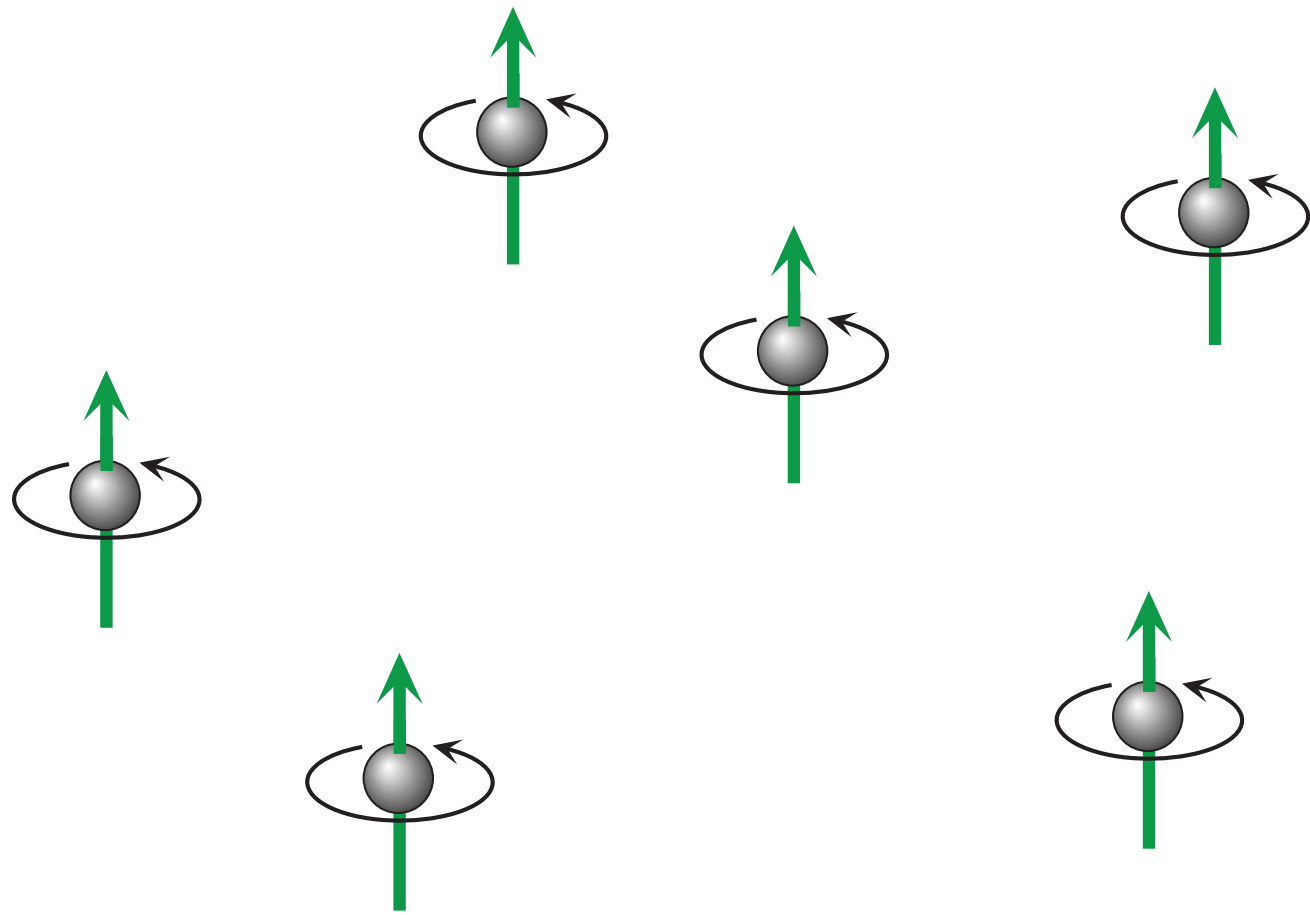


1900 1920 1940 1960 1980 2000

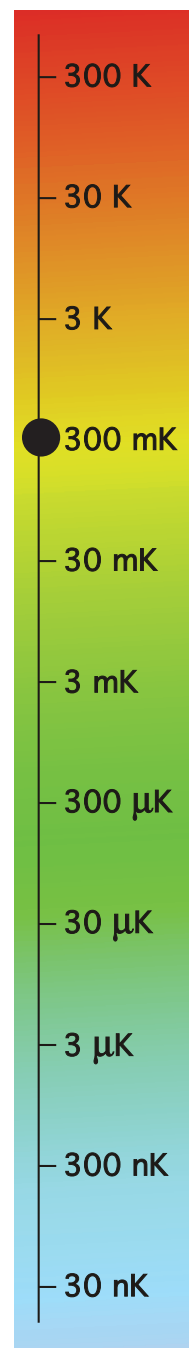
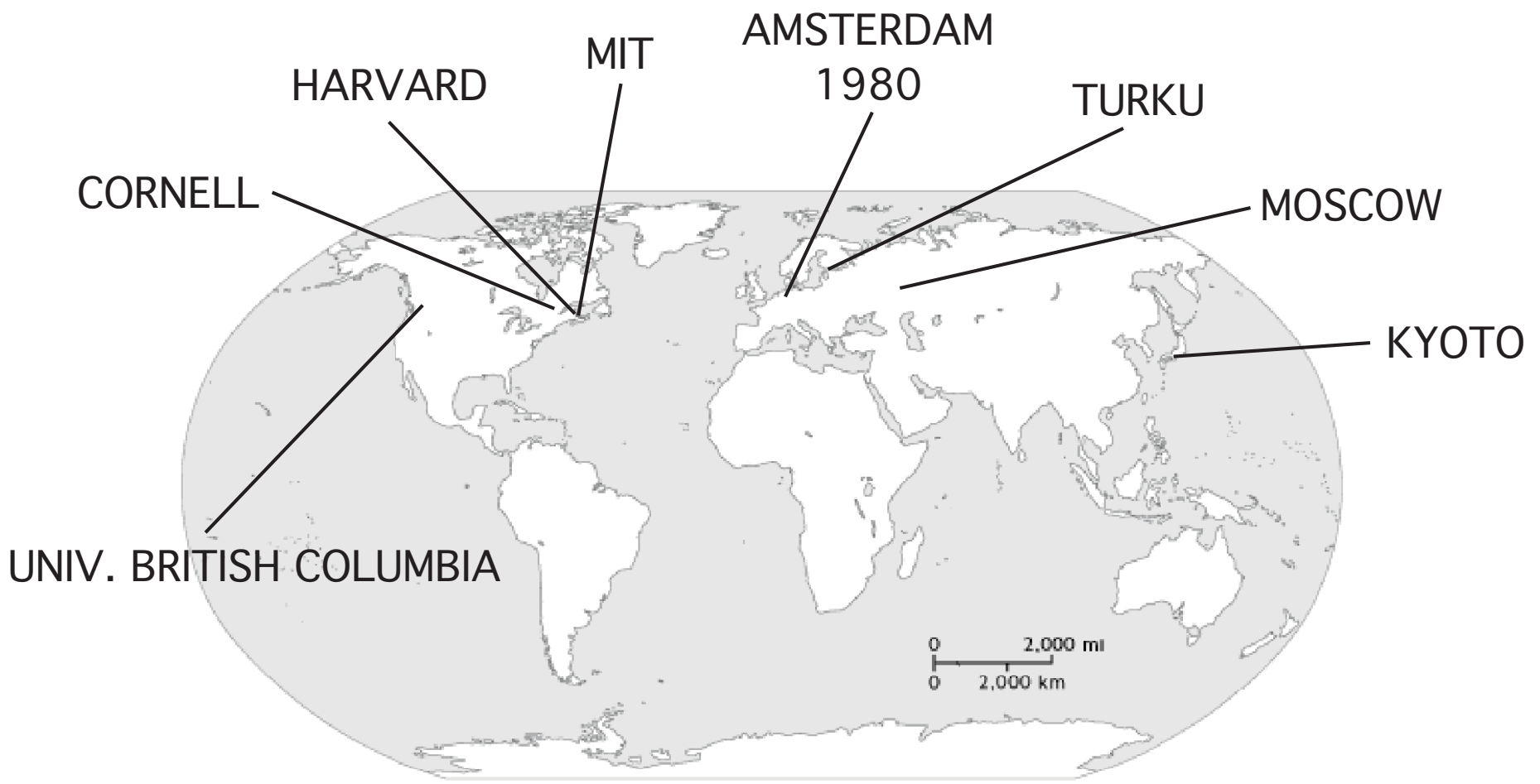
# THE SEARCH BEGINS FOR BOSE-EINSTEIN CONDENSATION IN A REAL GAS

See the article [Possible "New" Quantum Systems](#)  
In *Physics Review Letters*, Volume 36, Number 15 pages 910–913 (1976)

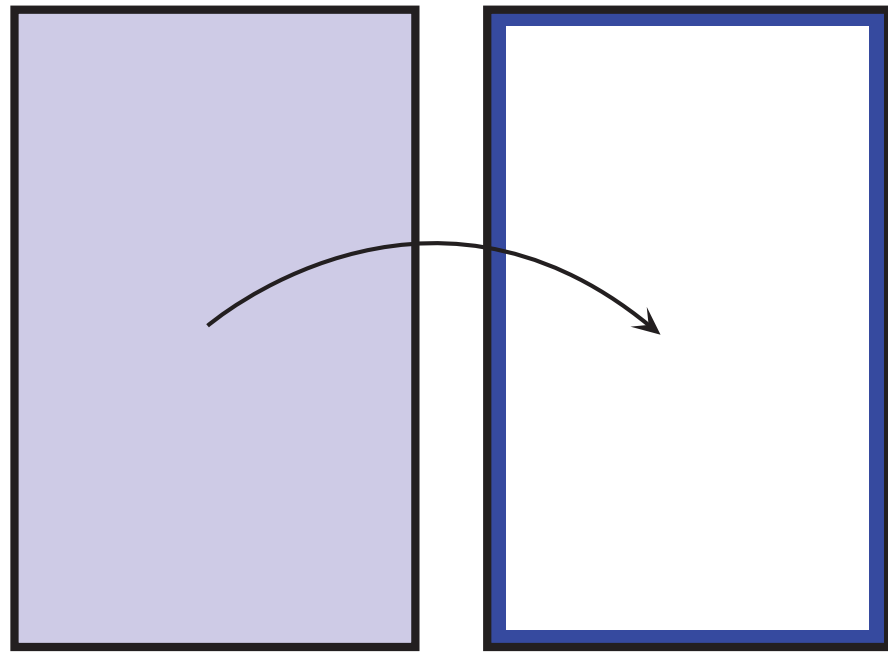




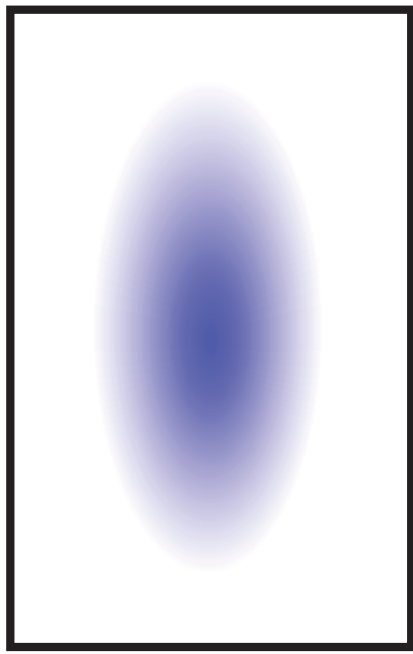
ATOMIC HYDROGEN WILL REMAIN A GAS DOWN TO ABSOLUTE ZERO IF ITS MAGNETIC MOMENTS ARE ALIGNED BY A MAGNETIC FIELD.



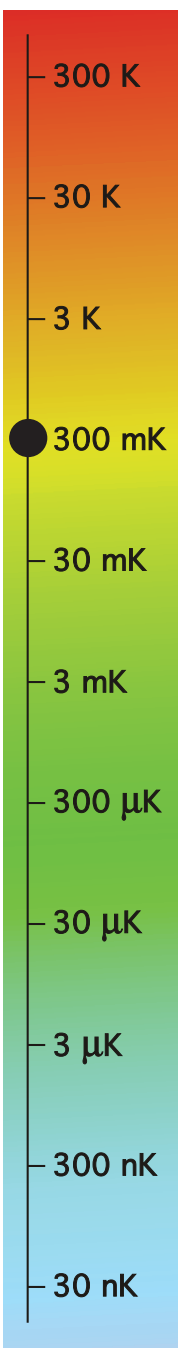
# GROUPS WORKING ON SPIN-POLARIZED ATOMIC HYDROGEN



CONFINEMENT BY WALLS

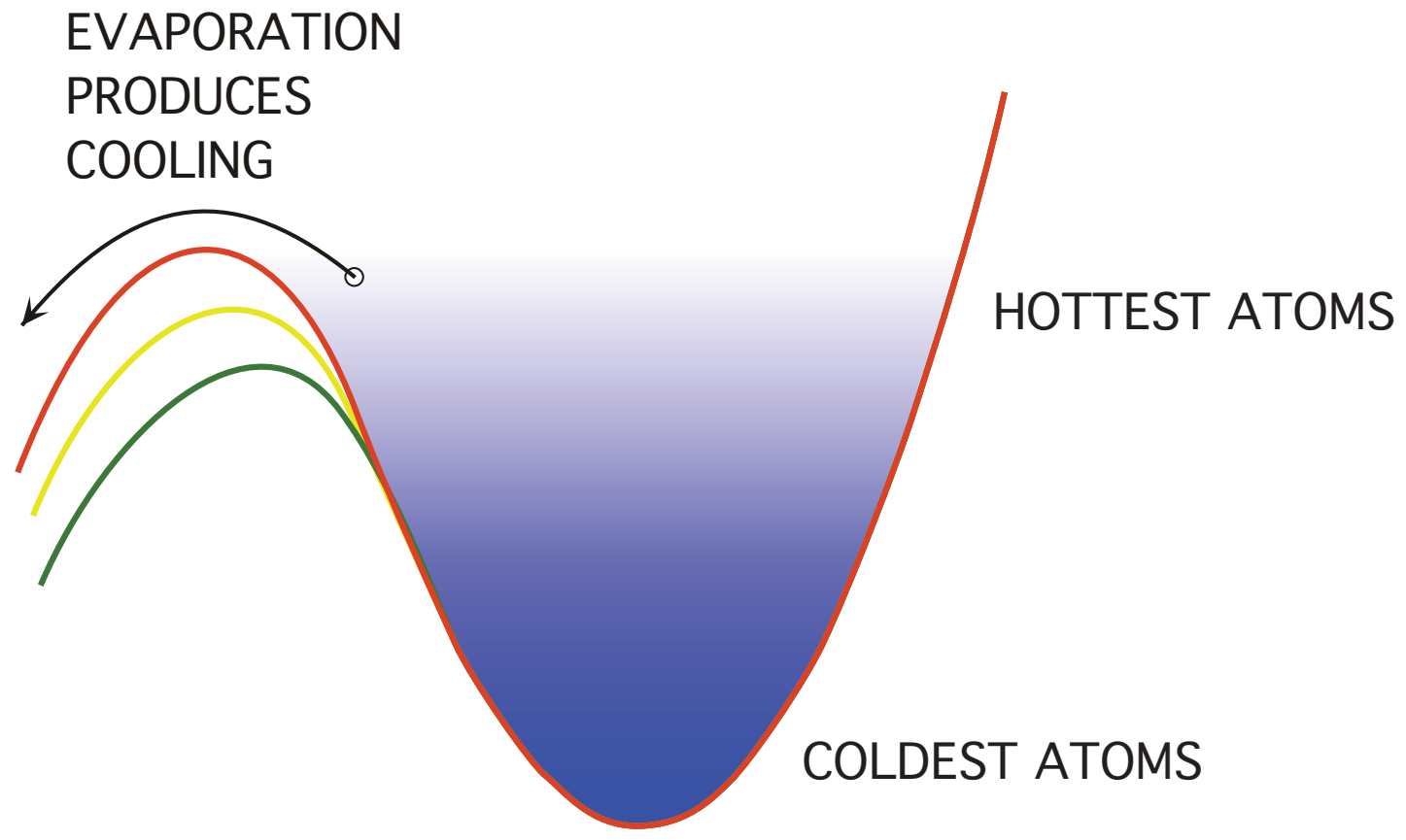


TRAPPING BY A MAGNETIC FIELD

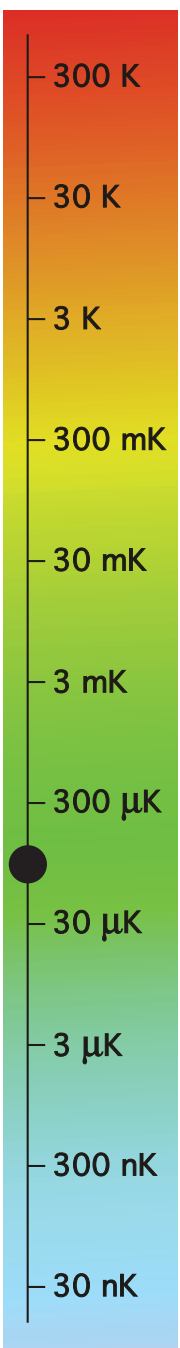
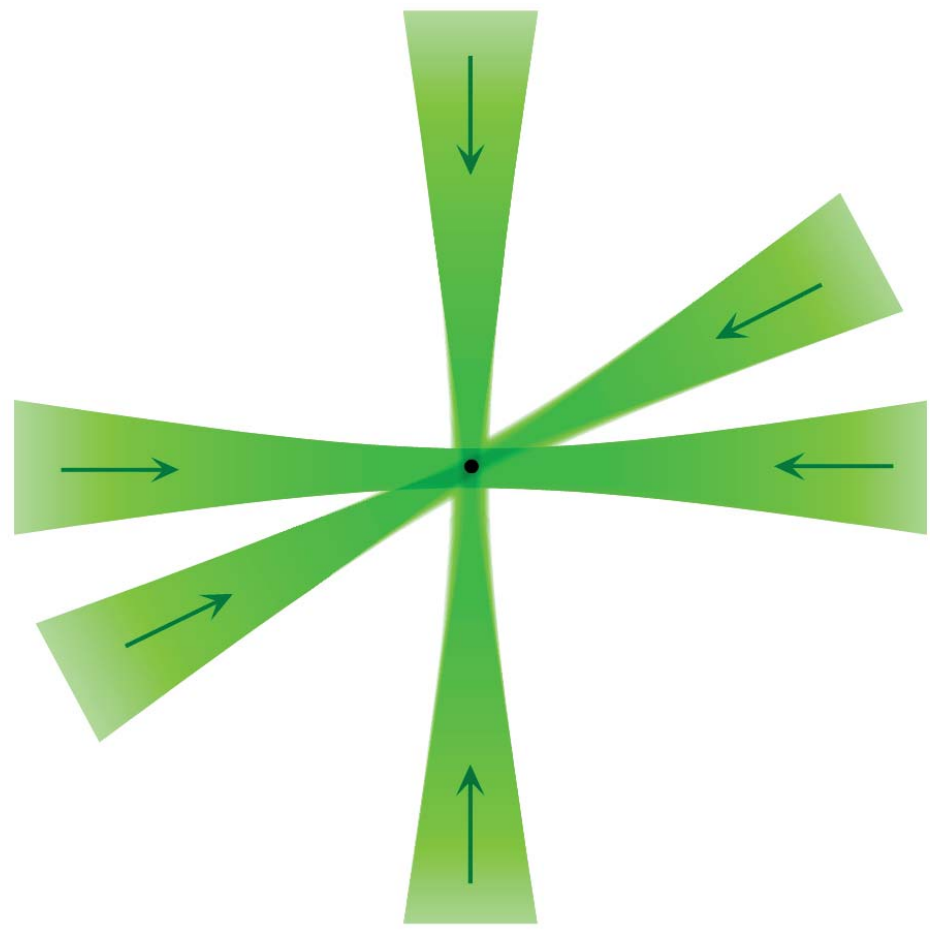


WALLS CAUSE THE MOMENTS TO FLIP; THEN THE ATOMS RECOMBINE INTO MOLECULES AND FREEZE OUT.

A MAGNETIC TRAP KEEPS THE ATOMS OFF THE WALLS.



IN 1986 HARALD HESS, A POSTDOCTORAL FELLOW IN MIT'S HYDROGEN GROUP, PROPOSES MAGNETIC TRAPPING AND EVAPORATIVE COOLING.



BEGINNING IN THE 1980s, METHODS WERE DEVELOPED TO COOL ATOMS INTO THE MICROKELVIN REGION OF TEMPERATURES USING LASERS.

1900

1920

1940

1960

1980

2000



## The Nobel Prize in Physics 1997

The Nobel Prize in Physics 1997 was awarded jointly to [Steven Chu](#), [Claude Cohen-Tannoudji](#) and [William D. Phillips](#) "for development of methods to cool and trap atoms with laser light".







1900

1920

1940

1960

1980

2000

LASER COOLING WORKS BEST WITH CERTAIN ATOMS SUCH AS LITHIUM (Li), SODIUM (Na), and RUBIDIUM (Rb).

BUT LASER COOLING ALONE CAN NOT GET THESE ATOMS COLD ENOUGH TO ACHIEVE BEC.

FOR THE FINAL STAGE OF COOLING ONE MUST TURN TO EVAPORATIVE COOLING.

THEN THE RACE BEGAN: LOWER THE TEMPERATURE WHILE INCREASING THE DENSITY.



300 K

30 K

3 K

300 mK

30 mK

3 mK

300  $\mu$ K

30  $\mu$ K

3  $\mu$ K

300 nK

30 nK



- ERIC CORNELL & CARL WIEMAN  
JILA (NIST AND UNIV. OF COLORADO)

Rb

- RANDALL HULET  
RICE UNIVERSITY

Li

- WOLFGANG KETTERLE  
MIT

Na

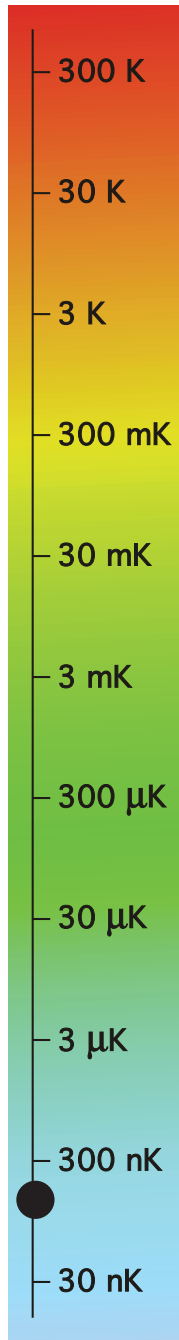
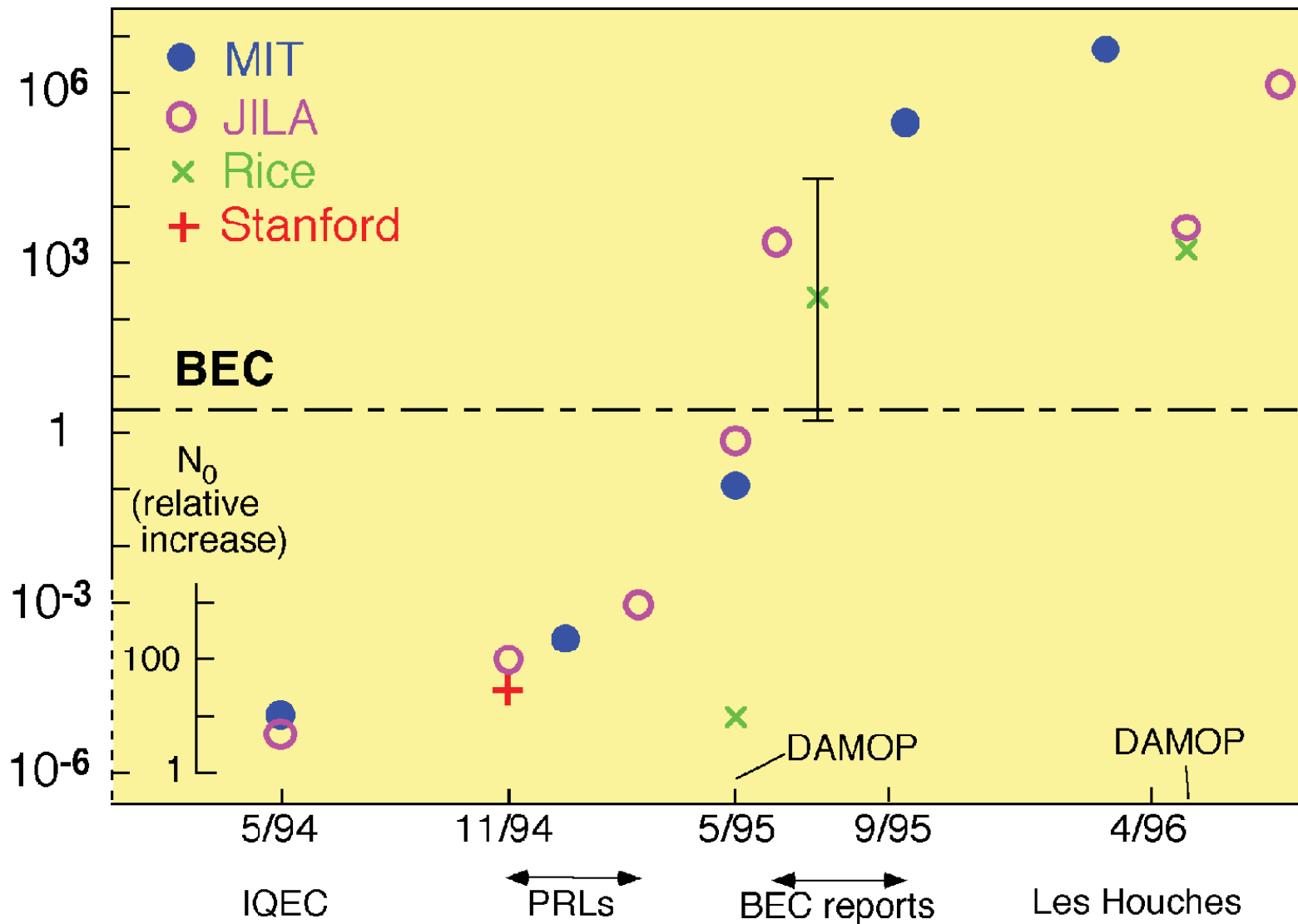




# Race to BEC

Evaporative cooling of alkali atoms

$N_0$  (Population of the lowest state) [or normalized density of atoms]



1900

1920

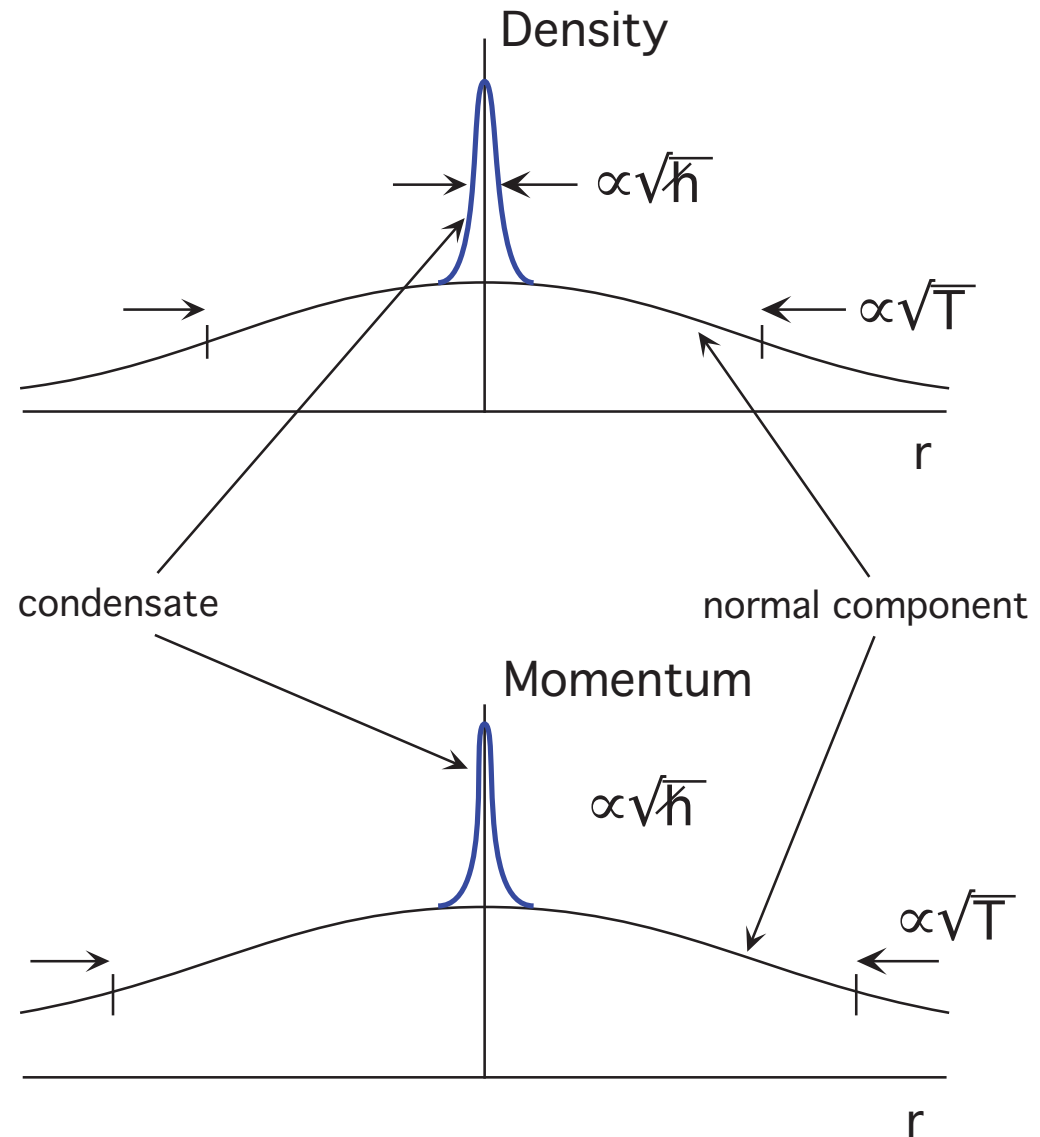
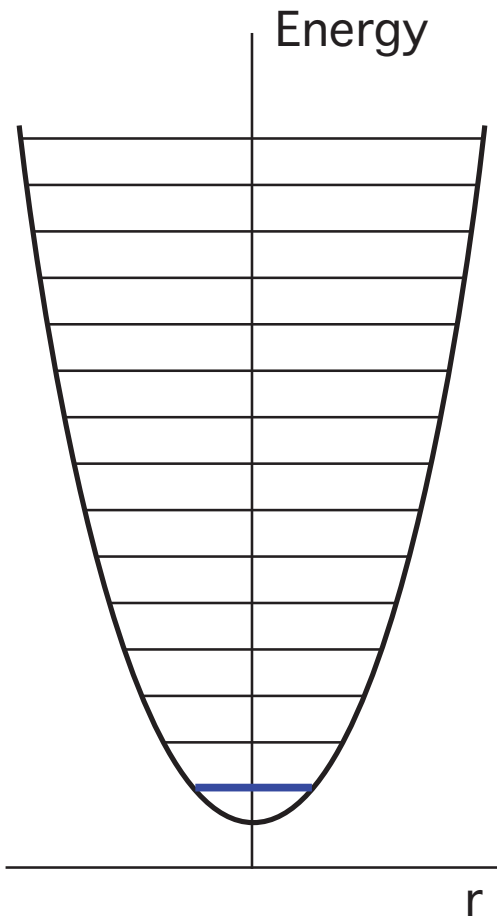
1940

1960

1980

2000

# Bose-Einstein Condensation in a Parabolic Trap



300 K

30 K

3 K

300 mK

30 mK

3 mK

300  $\mu$ K30  $\mu$ K3  $\mu$ K

300 nK

30 nK

1900

1920

1940

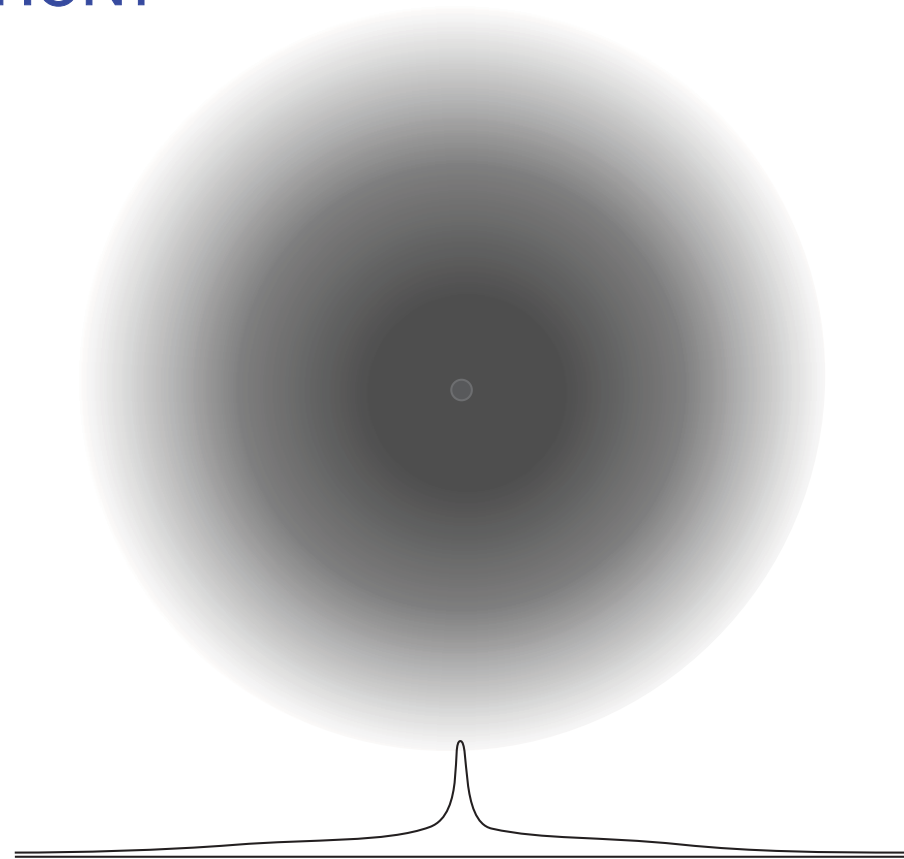
1960

1980

2000

## HOW DID THEY KNOW THEY HAD BOSE-EINSTEIN CONDENSATION?

IN THE TRAP, ATOMS IN THE CONDENSATE ARE ALMOST AT REST, THE REMAINDER HAVE THERMAL SPEEDS.



WHEN THE TRAP IS TURNED OFF THE THERMAL ATOMS SPEED AWAY, BUT THE CONDENSATE ATOMS REMAIN NEAR THE ORIGIN.

300 K

30 K

3 K

300 mK

30 mK

3 mK

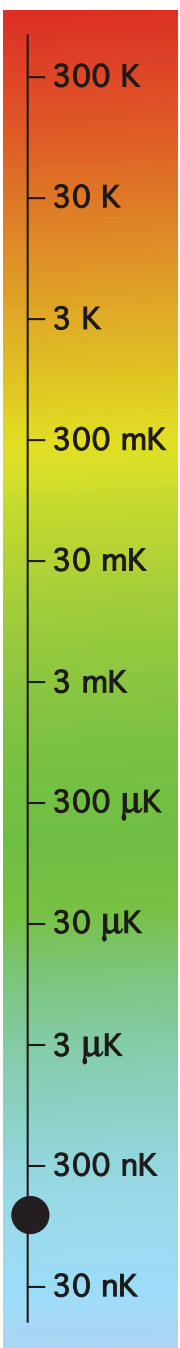
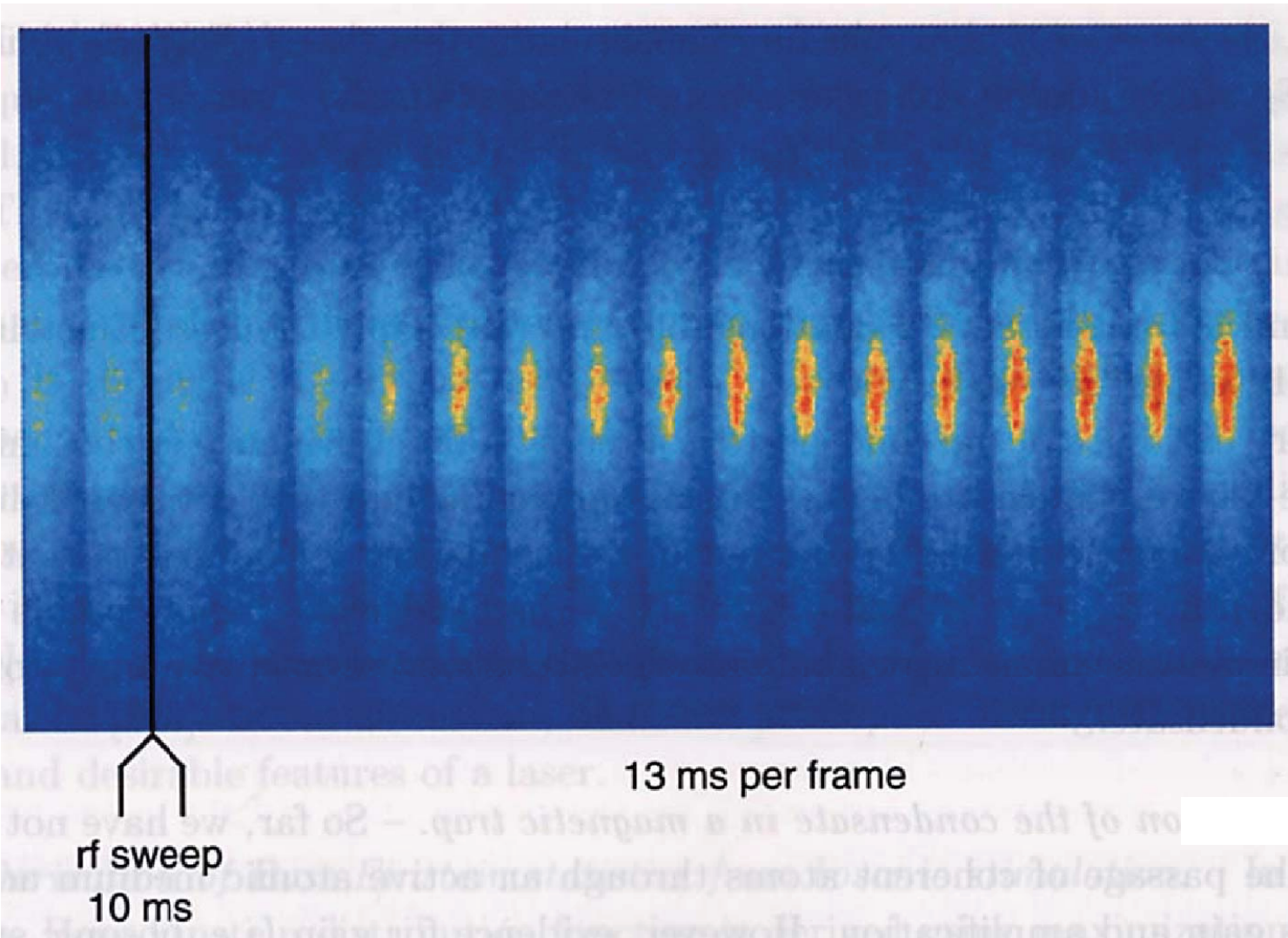
300 μK

30 μK

3 μK

300 nK

30 nK



SUCCESSIVE REAL SPACE IMAGES OF A SODIUM CONDENSATE FORMING IN A KETTERLE TRAP

1900

1920

1940

1960

1980

2000



## The Nobel Prize in Physics 2001

The Nobel Prize in Physics 2001 was awarded jointly to [Eric A. Cornell](#), [Wolfgang Ketterle](#) and [Carl E. Wieman](#) "for the achievement of Bose-Einstein condensation in dilute gases of alkali atoms, and for early fundamental studies of the properties of the condensates".

300 K

30 K

3 K

300 mK

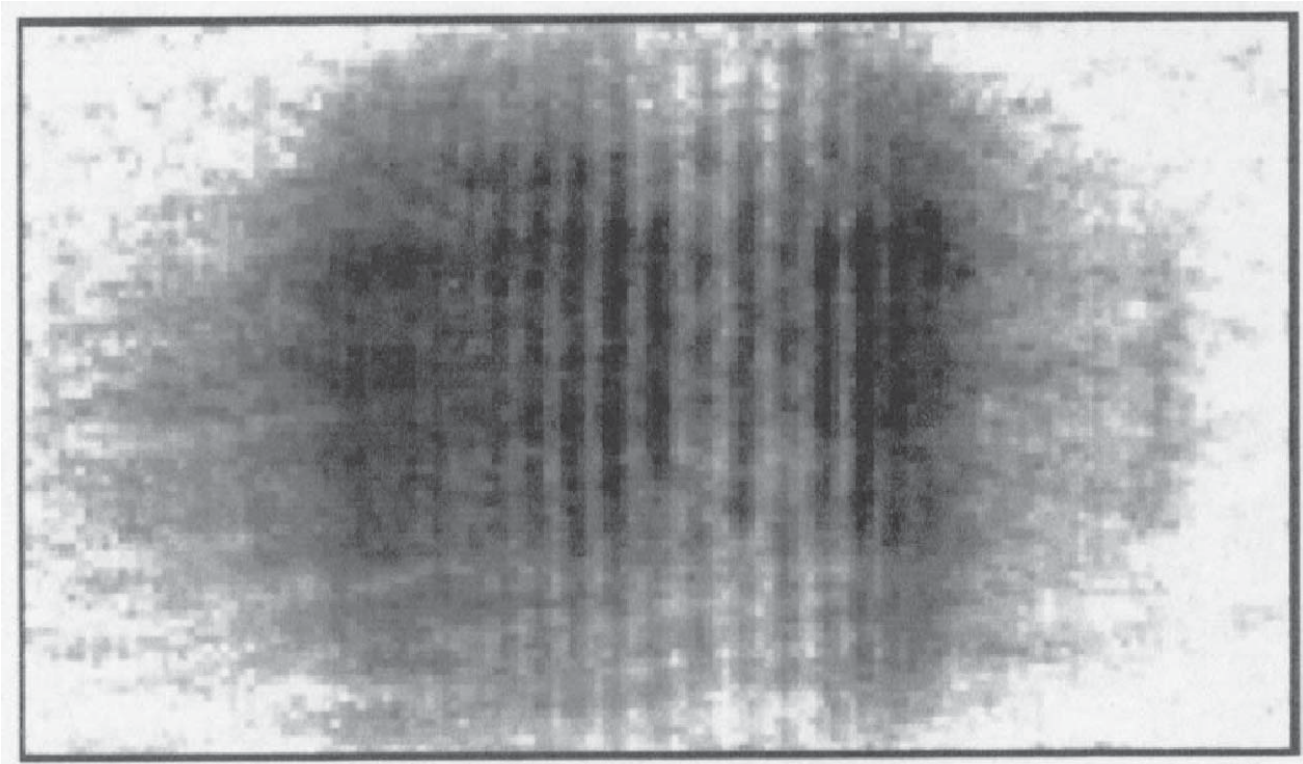
30 mK

3 mK

300  $\mu$ K30  $\mu$ K3  $\mu$ K

300 nK

30 nK



INTERFERENCE OF MATTER WAVES





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8.044 Statistical Physics I  
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