

**Retarded electromagnetic Green's function. Energy balance in radiation.**

**Reading:** Schwinger, Chaps. 31, 32.

**1. Charge moving at constant velocity.** (*Schwinger, Problems 1 & 2, Chap 31*)

(a) A particle with charge  $e$  moves with constant velocity  $\mathbf{v}$ . Its position is given by  $\mathbf{r}(t) = \mathbf{v}t$ . Construct the potentials,  $\phi$  and  $\mathbf{A}$ , in the Lorentz gauge, and show that

$$\phi = \frac{e}{\sqrt{(\mathbf{r} - \mathbf{v}t)^2 - \frac{v^2}{c^2} \mathbf{r}_\perp^2}}, \quad \mathbf{A} = \frac{\mathbf{v}}{c} \phi \quad (1)$$

where  $\mathbf{r}_\perp$  is the component of particle radius vector perpendicular to the velocity.

(b) What are the electric and magnetic fields for this particle?

(Hint: A particle in uniform motion is not very different from a particle at rest. One can either follow the route suggested in the textbook and use Eqs. (31.20),(31.21), or apply a Lorentz transformation to the fields of a stationary particle.)

**2. Potentials and fields of an arbitrarily moving charge.** (*Schwinger, Problems 5 & 6, Chap 31*)

(a) The charge and current densities of a point charge are given by  $\rho(\mathbf{r}, t) = e\delta(\mathbf{r} - \mathbf{r}(t))$ ,  $\mathbf{j}(\mathbf{r}, t) = e\mathbf{v}(t)\delta(\mathbf{r} - \mathbf{r}(t))$ . From the retarded potentials Eqs. (31.49),(31.50), derive the Lienard-Wiehart potentials

$$\phi(\mathbf{r}, t) = \frac{e}{|\mathbf{r} - \mathbf{r}(t')| - (\mathbf{r} - \mathbf{r}(t')) \cdot \frac{\mathbf{v}(t')}{c}}, \quad \mathbf{A}(\mathbf{r}, t) = \frac{\mathbf{v}(t')}{c} \phi(\mathbf{r}, t) \quad (2)$$

where the retarded time  $t'$  is defined by

$$t - t' = \frac{|\mathbf{r} - \mathbf{r}(t')|}{c} \quad (3)$$

(b) From potentials of part (a) compute  $\mathbf{E}$  and  $\mathbf{B}$ . Express the answer as the sum of a “velocity” part (involving  $\mathbf{v}$  only, and asymptotic to  $1/r^2$ ), and an “acceleration” part (proportional to  $\dot{\mathbf{v}}$  and asymptotic to  $1/r$ ). Only the latter is significant for radiation.

**3. Radiation of an electron moving in a magnetic field.** (*Schwinger, Problem 1, Chap 32*)

A nonrelativistic particle of charge  $e$  and mass  $m$  moves in a uniform magnetic field  $\mathbf{B}$ . Suppose the motion is confined to the plane perpendicular to  $\mathbf{B}$ . Calculate the radiated power  $P$  and express it in terms of particle energy,  $-P = dE/dt$ . For the energy  $E$ , derive an ordinary differential equation

$$-\frac{dE}{dt} = \gamma E \quad (4)$$

and find  $\gamma$  in term of  $\mathbf{B}$ . For an electron, find  $1/\gamma$  in seconds for a magnetic field of  $10^4$  gauss.

**4. Radiation of a classical hydrogen atom.** (*Schwinger, Problem 3, Chap 32*)

A nonrelativistic electron of charge  $e$  and mass  $m$  moves in a circular orbit under Coulomb forces produced by a proton. The average potential energy is related to the total energy by  $E = \frac{1}{2}\bar{V}$ .

Suppose that, as it radiates, the electron continues to move in a circle, and calculate the power radiated, and thereby  $-dE/dt$ , as a function of  $E$  (the relation is no longer linear). Integrate this result and find how long it takes for the energy to change from  $E_2$  to  $E_1$ .

Show that the electron reaches the center in a finite time. Calculate how long it takes an electron to hit the proton if it starts from an initial radius of  $r_0 = 10^{-8}$  cm. (This *radiation instability* of a classical atom was one of the reasons for the discovery of quantum mechanics.)

### 5. Dipole radiation.

(a) A nonrelativistic electron of charge  $e$  and mass  $m$  is driven by a time-dependent electric field  $\mathbf{E} = E_0 \hat{\mathbf{z}} \cos \omega t$ . Show that the radiation is given by that of an oscillating dipole  $\mathbf{d}(t) = e\mathbf{r}(t)$ , and find the angular distribution of radiated power *averaged over oscillation cycle*.

(b) A nonrelativistic electron of charge  $e$  and mass  $m$  moves in a uniform and constant magnetic field  $\mathbf{B}$ . As in part a), show that the radiation is given by that of a rotating dipole  $\mathbf{d}(t) = e\mathbf{r}(t)$ , and find the time-averaged radiated power angular distribution.