

## Lecture 26 (Dec. 11, 2017)

### 26.1 Harmonic Perturbations

We now study perturbations of the form

$$H_1(t) = Ve^{-i\omega t} + V^\dagger e^{i\omega t}. \quad (26.1)$$

Here,  $V$  is some operator acting on the Hilbert space that depends on the degrees of freedom in the system. We assume that  $V$  is “weak,” so that we can hope to treat it perturbatively.

If this perturbation acts for a long time, we expect that it will induce a transition from the initially prepared state  $|i\rangle$  to some final state  $|f\rangle$  where the system has absorbed energy  $\hbar\omega$  from the perturbation, i.e.,

$$E_f - E_i = \hbar\omega. \quad (26.2)$$

In the last class, we computed the transition amplitude for an arbitrary time-dependent perturbation at first order:

$$c_{fi}(t) = -\frac{i}{\hbar} \int_0^t dt' \langle f|H_1|i\rangle e^{i\omega_{fi}t'}. \quad (26.3)$$

For the case of harmonic perturbations, we find

$$\begin{aligned} c_{fi}(t) &= -\frac{i}{\hbar} \int_0^t dt' \left[ \langle f|V|i\rangle e^{i(\omega_{fi}-\omega)t'} + \langle f|V^\dagger|i\rangle e^{i(\omega_{fi}+\omega)t'} \right] \\ &= \frac{1}{\hbar} \left[ \langle f|V|i\rangle \left( \frac{1 - e^{i(\omega_{fi}-\omega)t}}{\omega_{fi} - \omega} \right) + \langle f|V^\dagger|i\rangle \left( \frac{1 - e^{i(\omega_{fi}+\omega)t}}{\omega_{fi} + \omega} \right) \right]. \end{aligned} \quad (26.4)$$

As  $t \rightarrow \infty$ ,  $|c_{fi}|^2$  is appreciable if either  $\omega_{fi} - \omega \approx 0$ , i.e.,  $E_f \approx E_i + \hbar\omega$  (absorption), or if  $\omega_{fi} + \omega \approx 0$ , i.e.,  $E_f = E_i - \hbar\omega$  (emission).

We now specialize to the case where  $\omega$  is tuned so that

$$E_f - E_i \approx \hbar\omega. \quad (26.5)$$

In this case, the second term in Eq. (26.4) is small compared to the first, so we can write

$$c_{fi}(t) \approx \frac{1}{\hbar} \langle f|V|i\rangle \left( \frac{1 - e^{i(\omega_{fi}-\omega)t}}{\omega_{fi} - \omega} \right). \quad (26.6)$$

The transition probability is then

$$\begin{aligned} P_{fi}(t) &= |c_{fi}(t)|^2 \\ &= \frac{1}{\hbar^2} |\langle f|V|i\rangle|^2 \frac{|1 - e^{i(\omega_{fi}-\omega)t}|^2}{(\omega_{fi} - \omega)^2} \\ &= \frac{1}{\hbar^2} |\langle f|V|i\rangle|^2 \frac{\sin^2\left(\frac{(\omega_{fi}-\omega)t}{2}\right)}{\left(\frac{(\omega_{fi}-\omega)t}{2}\right)^2} t^2. \end{aligned} \quad (26.7)$$

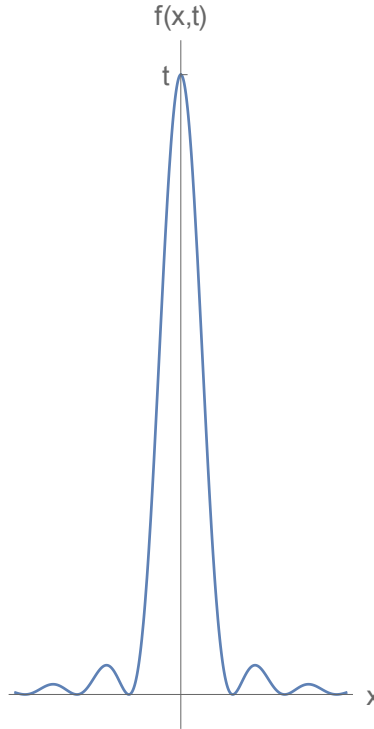
In this last line, we have simply used

$$\begin{aligned} |1 - e^{i\theta}|^2 &= \left| e^{i\theta/2} (e^{-i\theta/2} - e^{i\theta/2}) \right|^2 \\ &= \left| -2i \sin \frac{\theta}{2} \right|^2 \\ &= 4 \sin^2 \frac{\theta}{2}. \end{aligned} \quad (26.8)$$

Now, consider the function

$$f(x, t) = \frac{\sin^2 tx}{tx^2}, \quad (26.9)$$

which is shown below:



At  $x = 0$ , we have  $f(0, t) = t$ . As  $x \rightarrow \infty$ ,  $f(x, t) \rightarrow 0$ . The function oscillates with a period proportional to  $1/t$ . As  $t \rightarrow \infty$ , the function decays to zero very quickly, but  $f(0, t) = t \rightarrow \infty$ . Thus, this appears to be approaching something like a delta function. We can check this by noting that

$$\begin{aligned} \int_{-\infty}^{\infty} dx \frac{\sin^2 tx}{tx^2} &= \int_{-\infty}^{\infty} \frac{d(tx)}{t} \frac{\sin^2 tx}{tx^2} \\ &= \int_{-\infty}^{\infty} du \frac{\sin^2 u}{u^2} \\ &= \pi, \end{aligned} \quad (26.10)$$

which we see is independent of  $t$ . We conclude that

$$\lim_{t \rightarrow \infty} \frac{\sin^2 tx}{tx^2} = \pi \delta(x). \quad (26.11)$$

Thus,

$$\lim_{t \rightarrow \infty} P_{fi}(t) = \frac{1}{\hbar^2} |\langle f|V|i \rangle|^2 \pi \delta\left(\frac{\omega_{fi} - \omega}{2}\right) t = \frac{2\pi}{\hbar^2} |\langle f|V|i \rangle|^2 \delta(\omega_{fi} - \omega) t. \quad (26.12)$$

This may seem odd, because the result we found is that as  $t \rightarrow \infty$ , the probability of transition diverges to infinity, because it is proportional to  $t$ . The interpretation here is that there is a constant transition rate,

$$R_{fi} = \frac{P_{fi}}{t} = \frac{2\pi}{\hbar^2} |\langle f|V|i \rangle|^2 \delta(\omega_{fi} - \omega). \quad (26.13)$$

In general, there will be several possible final states  $|f\rangle$  to which  $H_1$  can induce a transition. The total transition rate is then given by the sum over all final states  $|f\rangle$  of this rate. We define the density of states  $\rho(E)$  such that  $\rho(E) dE$  is the number of states between  $E$  and  $E + dE$ . Then the net transition rate out of the initial state is given by

$$\begin{aligned} W_{i,\text{out}} &= \sum_f R_{fi} \\ &= \int dE_f \rho(E_f) \frac{2\pi}{\hbar^2} |\langle f|V|i\rangle|^2 \delta(\omega_f - \omega_i - \omega). \end{aligned} \quad (26.14)$$

Changing variables in the delta function, we reach

$$W_{i,\text{out}} = \int dE_f \rho(E_f) \frac{2\pi}{\hbar} |\langle f|V|i\rangle|^2 \delta(E_f - E_i - \hbar\omega). \quad (26.15)$$

This famous result is known as *Fermi's Golden Rule*.

### 26.1.1 The Photoelectric Effect

Consider an electron in a hydrogen atom interacting with an external EM field. This system has Hamiltonian

$$H = \frac{(\mathbf{p} - \frac{e}{c}\mathbf{A})^2}{2m} - \frac{e^2}{r}, \quad (26.16)$$

where  $\mathbf{A}$  is the external vector potential. Take the vector potential to be of the form

$$\mathbf{A} = \mathbf{A}_0 \cos(\omega t - \mathbf{k} \cdot \mathbf{r}), \quad (26.17)$$

and assume that  $\mathbf{A}_0$  is weak. We can then write

$$H = \underbrace{\frac{\mathbf{p}^2}{2m} - \frac{e^2}{r}}_{H_0} - \frac{e}{2mc} (\mathbf{p} \cdot \mathbf{A} + \mathbf{A} \cdot \mathbf{p}) + \frac{e^2}{2mc^2} \mathbf{A}^2. \quad (26.18)$$

where  $H_0$  is the hydrogen atom Hamiltonian. The final term, proportional to  $\mathbf{A}^2$ , does not contribute to the problem we are considering at first order in perturbation theory, so we will simply drop it here.

It is convenient to work in Coulomb gauge,  $\nabla \cdot \mathbf{A} = 0$ . In this gauge,

$$(\mathbf{p} \cdot \mathbf{A} + \mathbf{A} \cdot \mathbf{p})\psi = (-i\hbar\nabla \cdot \mathbf{A} + 2\mathbf{A} \cdot \mathbf{p})\psi = 2\mathbf{A} \cdot \mathbf{p}\psi. \quad (26.19)$$

Thus, we can use

$$H = H_0 + H_1 \quad (26.20)$$

with

$$\begin{aligned} H_1 &= -\frac{e}{mc} \cos(\omega t - \mathbf{k} \cdot \mathbf{r}) \mathbf{A}_0 \cdot \mathbf{p} \\ &= -\frac{e}{2mc} \left( e^{i(\omega t - \mathbf{k} \cdot \mathbf{r})} + e^{-i(\omega t - \mathbf{k} \cdot \mathbf{r})} \right) \mathbf{A}_0 \cdot \mathbf{p}. \end{aligned} \quad (26.21)$$

The first term gives the transition rate proportional to  $\delta(E_f - E_i + \hbar\omega)$ , while the second term gives the transition rate proportional to  $\delta(E_f - E_i - \hbar\omega)$ .

If we take the atom to initially be in its ground state, then only the second term can contribute, so Fermi's Golden Rule gives

$$R_{0 \rightarrow f} = \frac{2\pi}{\hbar} \sum_{\mathbf{k}_f} \left| \langle \mathbf{k}_f | \frac{e}{2mc} \mathbf{A}_0 \cdot \mathbf{p} e^{-i\mathbf{k} \cdot \mathbf{r}} | 0 \rangle \right|^2 \delta(E_f - E_0 - \hbar\omega). \quad (26.22)$$

Here,  $|0\rangle$  denotes the ground state of the hydrogen atom, with wavefunction

$$\psi_0(\mathbf{x}) = \langle \mathbf{x} | 0 \rangle = \frac{1}{(\pi a_0^3)^{1/2}} e^{-r/a_0}, \quad (26.23)$$

where  $a_0$  is the Bohr radius. The ket  $|\mathbf{k}_f\rangle$  is a plane wave state with wave vector  $\mathbf{k}_f$ , which describes the electron that is kicked out of the atom by the perturbation.

We can compute the matrix element in the transition rate by putting the system in a box of side length  $L$ , which gives

$$\begin{aligned} \frac{e}{2mc} \mathbf{A}_0 \cdot \int d^3x \frac{e^{-i\mathbf{k}_f \cdot \mathbf{r}}}{L^{3/2}} (-i\hbar \nabla) \frac{e^{-r/a_0}}{(\pi a_0^3)^{1/2}} &= \frac{e}{2mc} \mathbf{A}_0 \cdot \int \frac{d^3x}{L^{3/2}} \left( i\hbar \nabla \left( e^{-i\mathbf{k}_f \cdot \mathbf{r}} \right) \right) \frac{e^{-r/a_0}}{(\pi a_0^3)^{1/2}} \\ &= \frac{e\hbar}{2mc} \frac{\mathbf{A}_0 \cdot \mathbf{k}_f}{L^{3/2} (\pi a_0^3)^{1/2}} \int d^3x e^{-i\mathbf{k}_f \cdot \mathbf{r}} e^{-r/a_0} \\ &= \frac{e\hbar}{2mc} \frac{\mathbf{A}_0 \cdot \mathbf{k}_f}{L^{3/2} (\pi a_0^3)^{1/2}} \frac{8\pi}{a_0 (k_f^2 + 1/a_0^2)^2}. \end{aligned} \quad (26.24)$$

Note that we have ignored the factor of  $e^{i\mathbf{k} \cdot \mathbf{r}}$  in this matrix element; this was not an accident. For typical wavelengths involved in photo-ionization,  $|\mathbf{k}| \ll |\mathbf{k}_f|$ .

For the sum over final states, we make the replacement

$$\sum_{\mathbf{k}_f} \rightarrow L^3 \int \frac{d^3k_f}{(2\pi)^3}. \quad (26.25)$$

The logic is that, in a box of size  $L$ , the set of allowed momenta has spacing  $2\pi/L$ , and so each differential element in the integral is of the form

$$\frac{dk_x}{(2\pi/L)}, \quad (26.26)$$

which gives the form of the replacement above.

Putting everything together, we then have

$$\begin{aligned} R_{0 \rightarrow f} &= \frac{2\pi}{\hbar} \frac{L^3}{L^3} \int \frac{d^3k_f}{(2\pi)^3} \left( \frac{e\hbar}{2mc} \right)^2 \frac{(\mathbf{A}_0 \cdot \mathbf{k}_f)^2}{\pi a_0^3} \frac{(8\pi)^2}{a_0^2 (k_f^2 + 1/a_0^2)^4} \delta(E_f - E_0 - \hbar\omega) \\ &= \frac{e^2 \hbar}{2m^2 c^2 a_0^5} \int \frac{dk_f k_f^2 d\Omega}{(2\pi)^3} \frac{(\mathbf{A}_0 \cdot \mathbf{k}_f)^2 (8\pi)^2}{(k_f^2 + 1/a_0^2)^4} \delta\left( \frac{\hbar^2 k_f^2}{2m} - E_0 - \hbar\omega \right) \\ &= \frac{e^2 k_f}{2mc^2 a_0^5 \hbar} \int \frac{d\Omega}{(2\pi)^3} \frac{(\mathbf{A}_0 \cdot \mathbf{k}_f)^2 (8\pi)^2}{(k_f^2 + 1/a_0^2)^4}, \end{aligned} \quad (26.27)$$

with  $k_f$  determined by energy conservation. The only angular dependence is in  $\mathbf{A}_0 \cdot \mathbf{k}_f$ , so using

$$\int d\Omega \cos^2 \theta = \frac{4\pi}{3} \quad (26.28)$$

we reach the final result,

$$R_{0 \rightarrow f} = \frac{16e^2 k_f^3 A_0^2}{3mc^2 \hbar a_0^5} \frac{1}{\left(k_f^2 + 1/a_0^2\right)^4}. \quad (26.29)$$

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